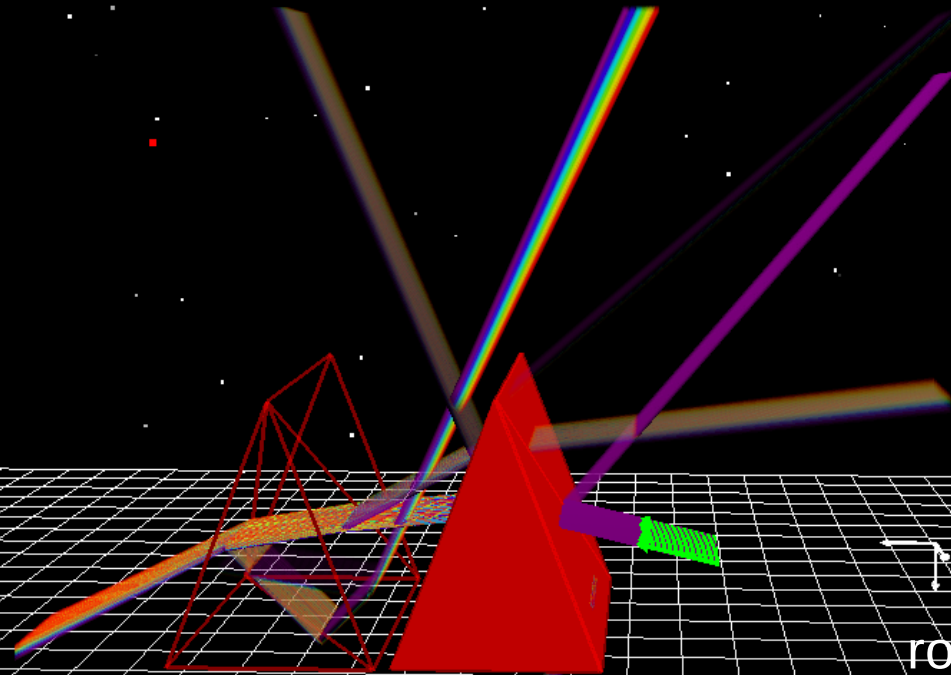
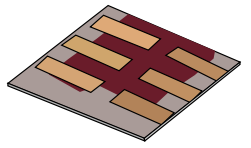


# Simulating light in opto-electronic devices, solar cells, sensors, and bio-sensors using OghmaNano



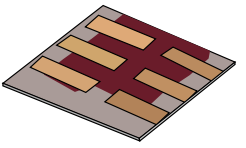
[roderick.mackenzie@oghma-nano.com](mailto:roderick.mackenzie@oghma-nano.com)

# Outline of the talk



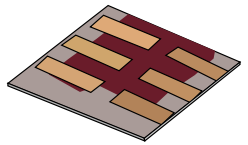
- In this talk we will cover:
  - What are optical simulations?
  - Why perform optical simulations?
  - What you need for accurate optical simulations
    - » Optical spectra
    - » Refractive index data ( $n$ )
    - » Optical absorption data ( $k$ )
      - The materials database
      - Importing  $n/k$  data into the model.
  - Setting up device structures
  - Running optical simulations using OghmaNano.
  - Light sources
  - Output files
  - Optical filter design task
  - Summary

# Download the software:



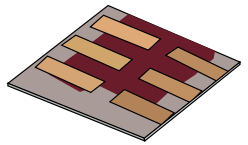
- Download all the software used in this talk from:
  - <http://www.oghma-nano.com/download.php>
- Please report bugs to:
  - [roderick.mackenzie@oghma-nano.com](mailto:roderick.mackenzie@oghma-nano.com)

# Outline of the talk

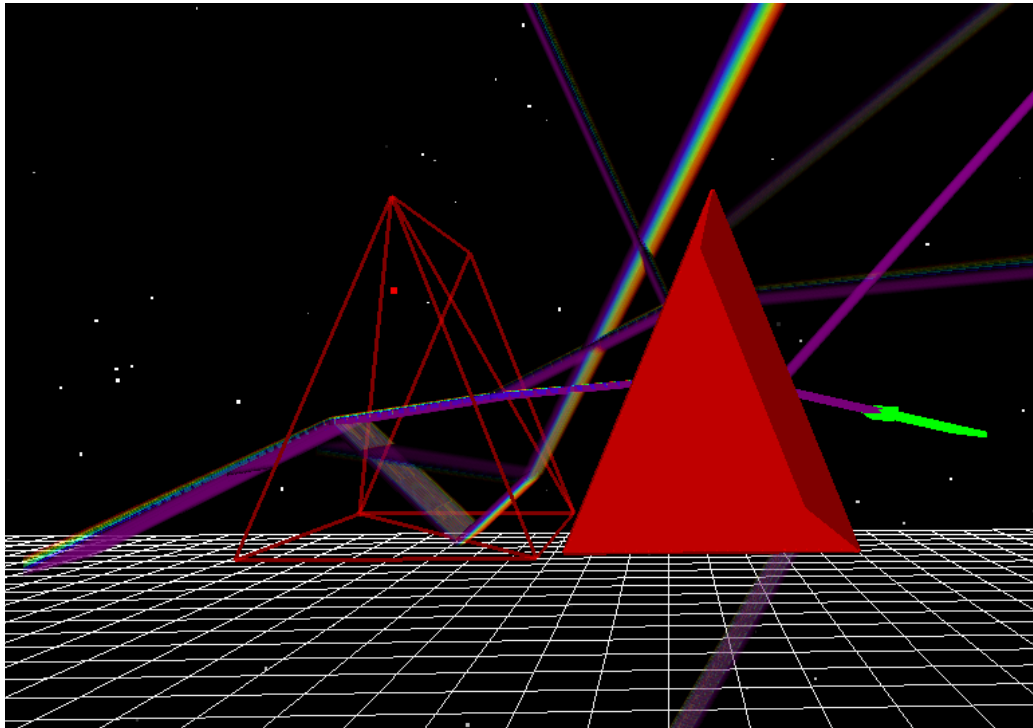


- In this talk we will cover:
  - **What are optical simulations?**
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  - What you need for accurate optical simulations
    - » Optical spectra
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# What are optical simulations?

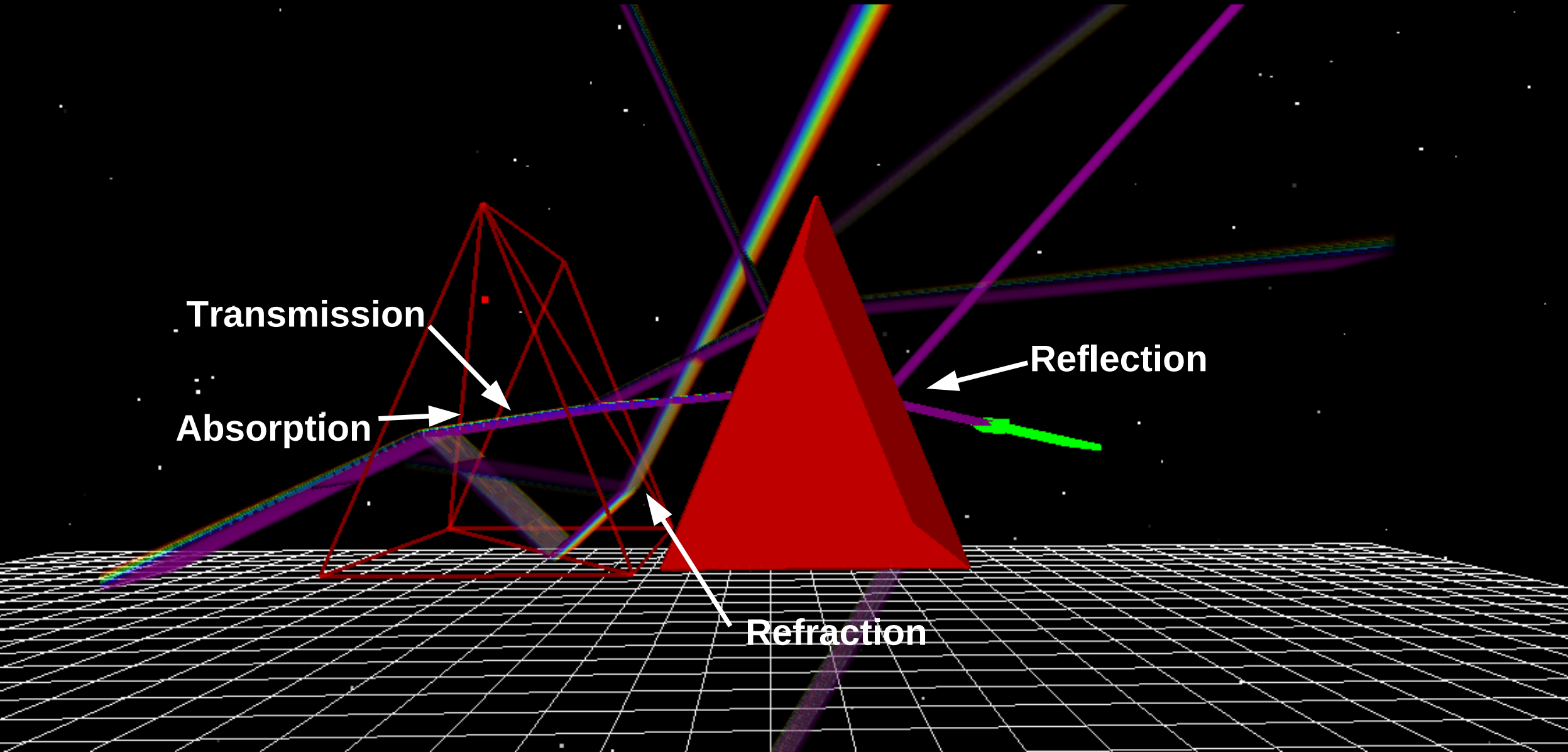


- Optical simulations enable you to understand how light interacts with your device.
- The front cover of this slide deck was an example of a simple optical simulation it showed light interacting with two prisms

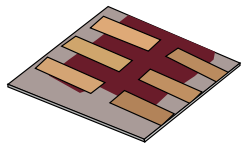


- You are able to predict
  - **Reflections**
  - **Transmission**
  - **Absorption**
  - **Refraction**
- And understand what this means for your device performance
  - What ever type of device it is.

Let's look at this picture in a bit more detail

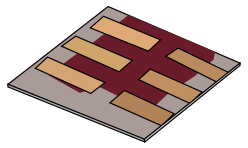


# Outline of the talk



- In this talk we will cover:
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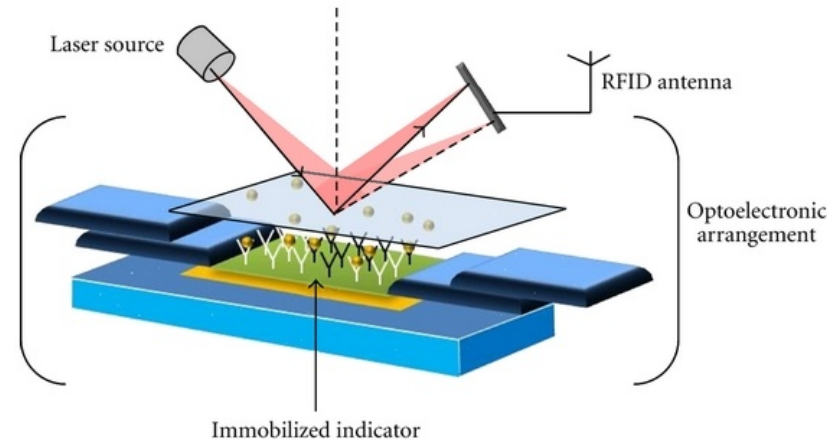
# Why perform optical simulations?



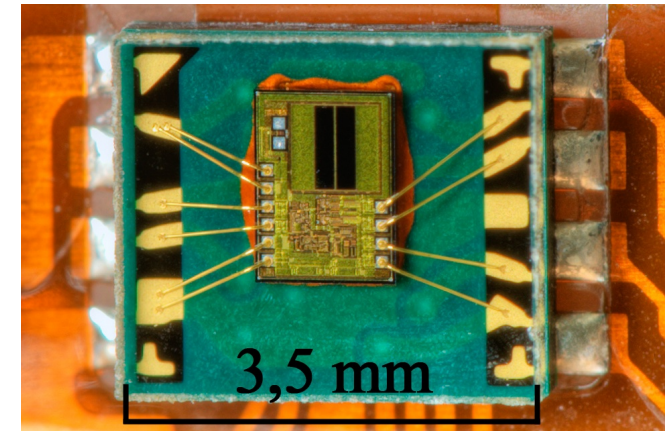
- Understanding how light interacts with devices is important for designing many classes of device including:



Solar cells



Bio-sensors



Optical sensors, DVD or telecommunications

- If you have photons in your device you will want to perform optical simulations at some point.

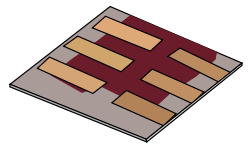
<https://doi.org/10.1155/2011/348218>

[https://commons.wikimedia.org/wiki/File:Photodetector\\_\(dvd\\_drive\).jpg](https://commons.wikimedia.org/wiki/File:Photodetector_(dvd_drive).jpg)

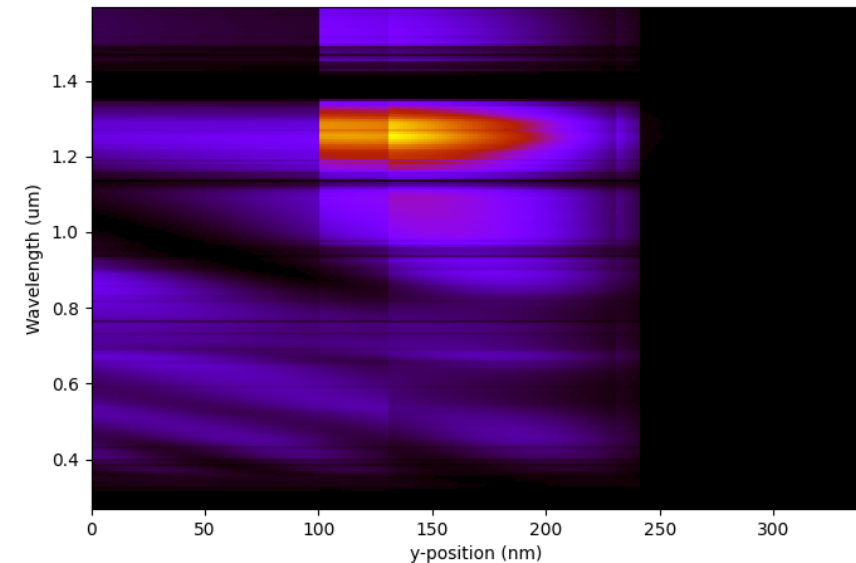
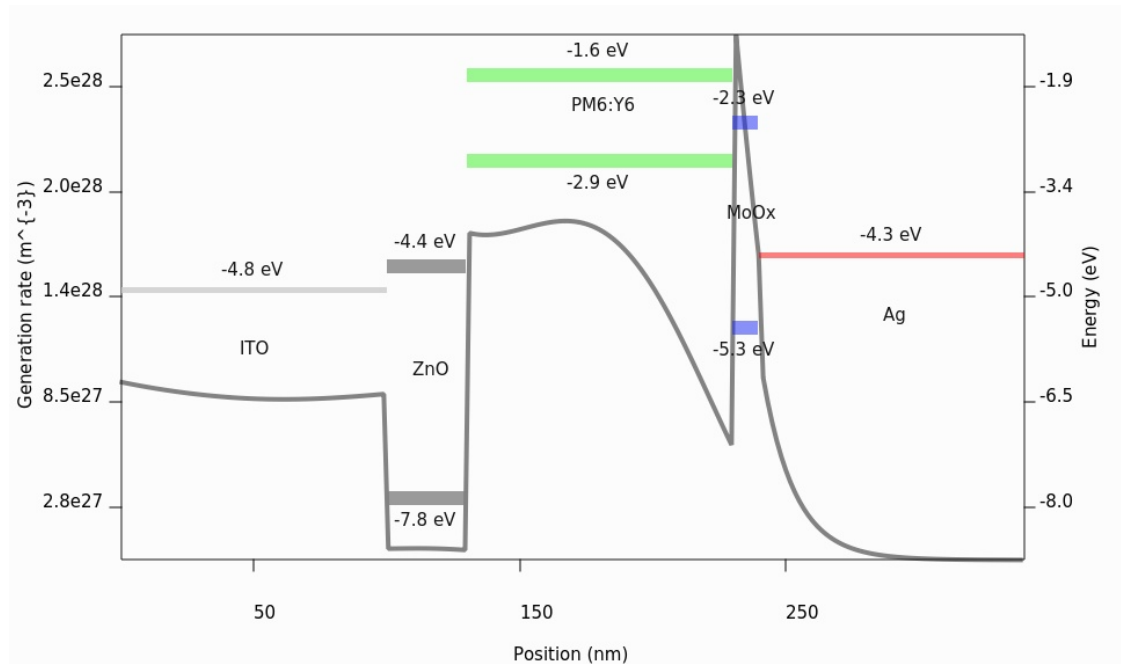


# Why perform optical simulations?

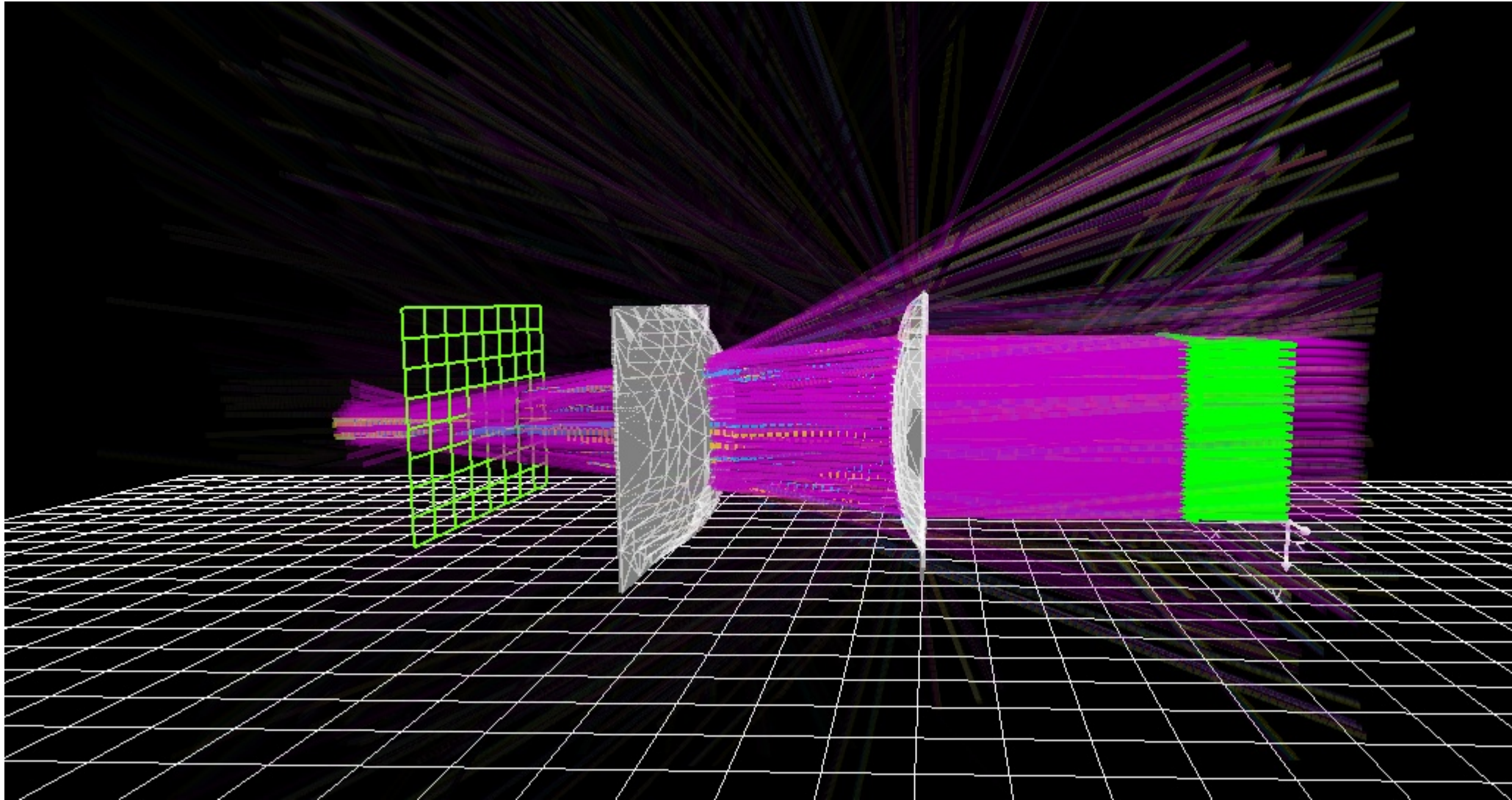
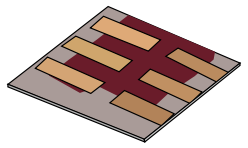
## Solar cell example:



- Optical simulations can tell you:
  - Where photons are being absorbed
  - At what wavelength
  - If your device is as efficient as it could be.

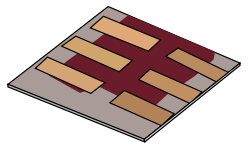


# Why perform optical simulations? Designing optical systems



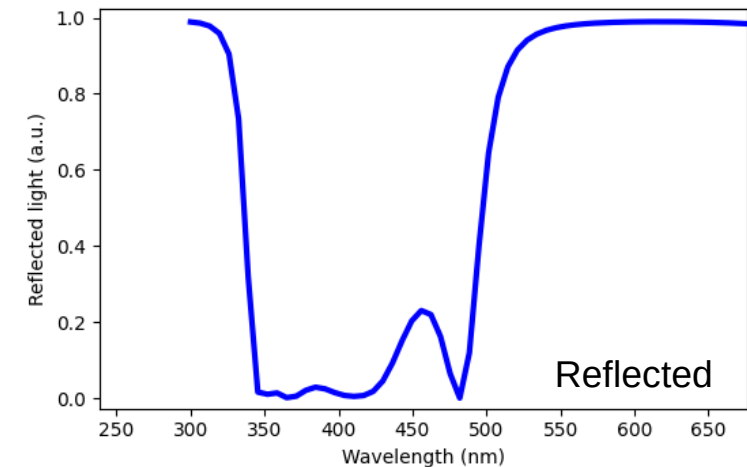
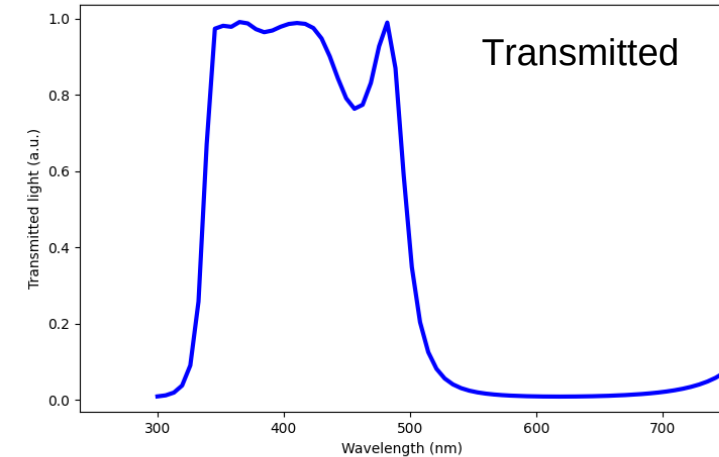
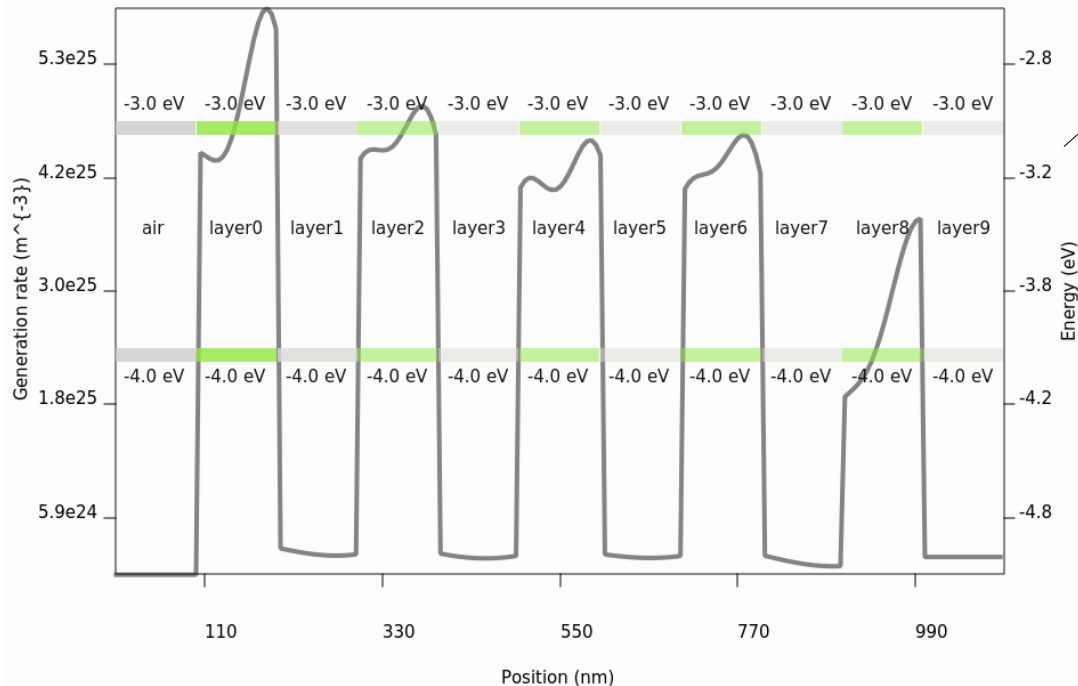
# Why perform optical simulations?

## Optical filter example:



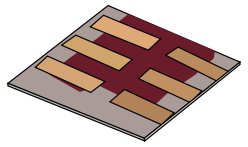
• Optical simulations allow you to:

- Play with the layer thickness before fabrication
- Play with the material properties before fabrication.



# Why perform optical simulations?

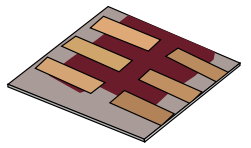
## They are easy to run and produce very reliable results



- All you need is the:
  - ***Optical spectra*** of the incident light
  - The ***refractive index*** of the material as a function of wavelength
  - The ***absorption of the material*** as a function of wavelength
  - And your ***device structure***

If you have this information your simulations will be 100% accurate.

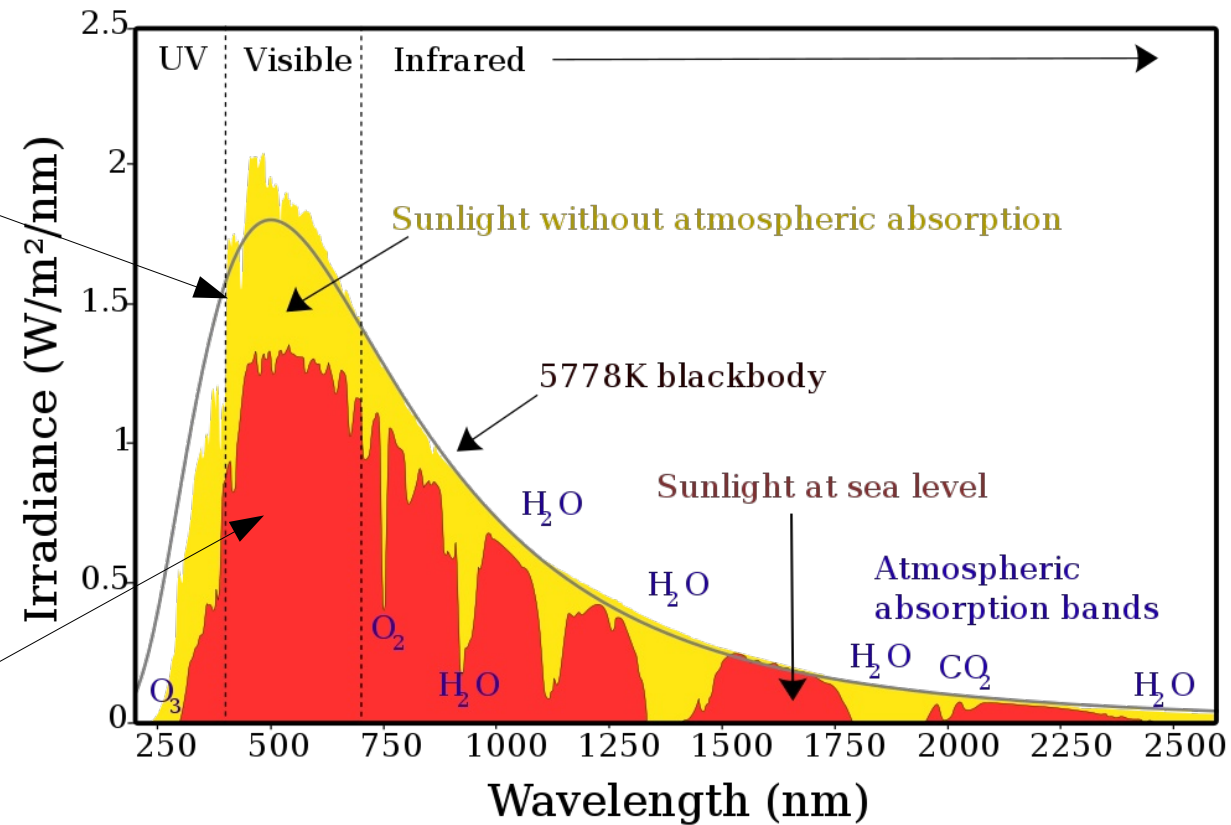
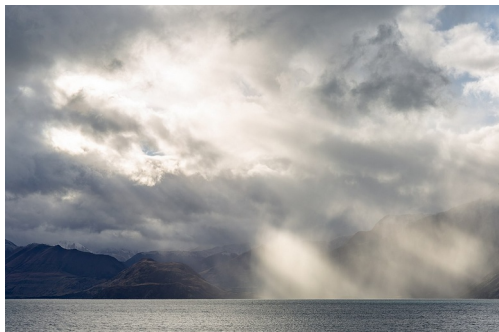
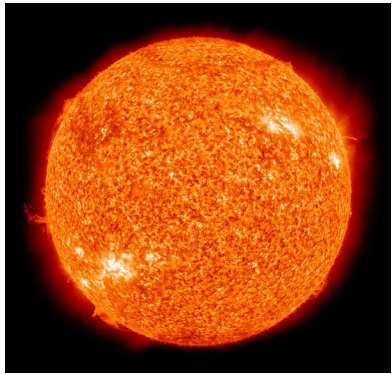
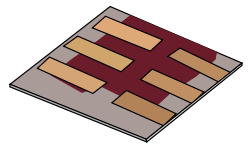
# Outline of the talk



- In this talk we will cover:
  - What are optical simulations?
  - Why perform optical simulations?
  - **What you need for accurate optical simulations**
    - » **Optical spectra**
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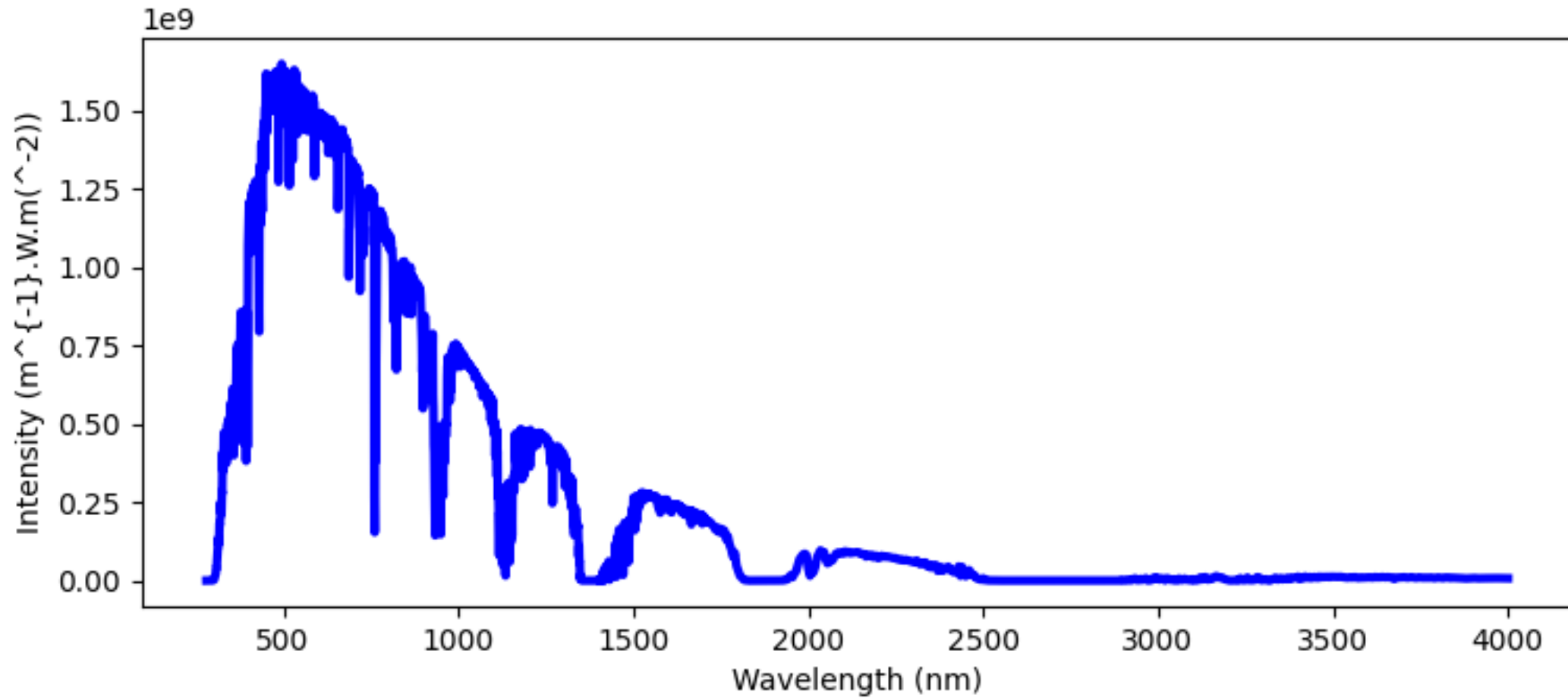
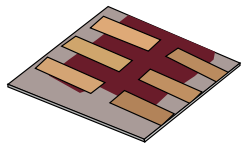


# The optical spectra of the Sun AM0, AM1.5G

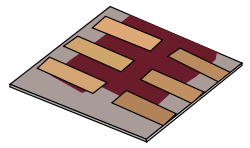


These spectra are known quantities.

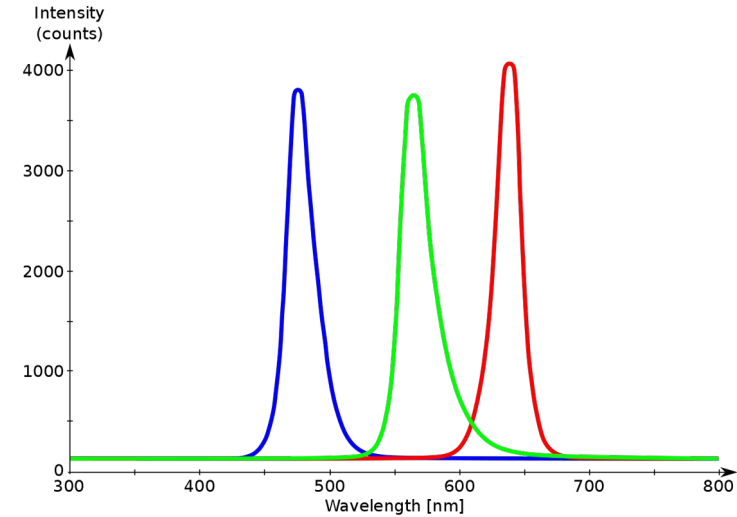
# AM1.5G in more detail



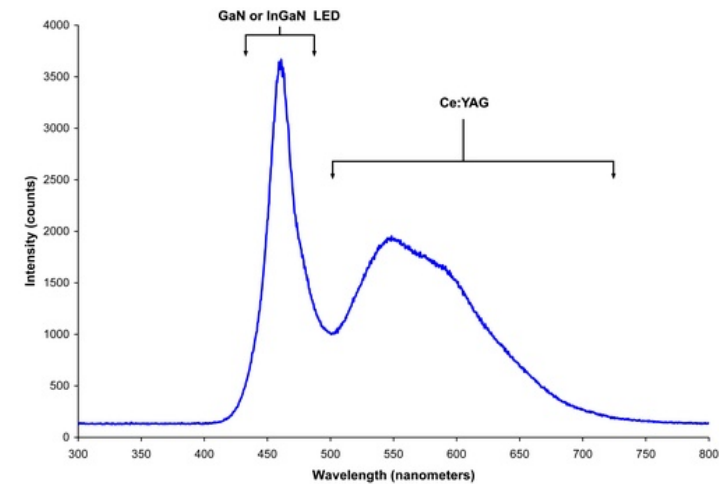
# Other examples of spectra commonly used in science/Engineering:



- Blue, Green and Red LEDs.



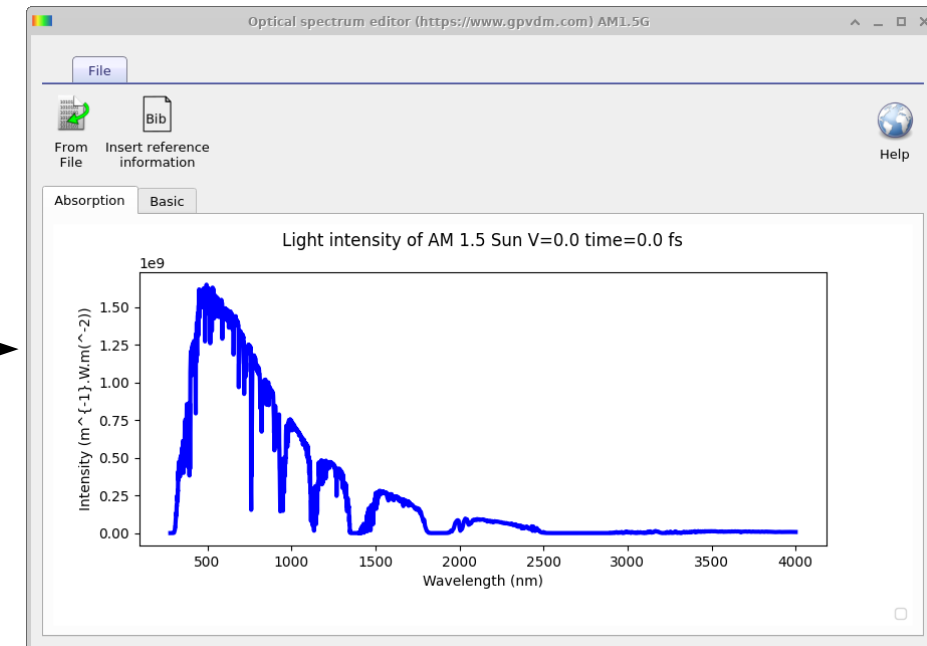
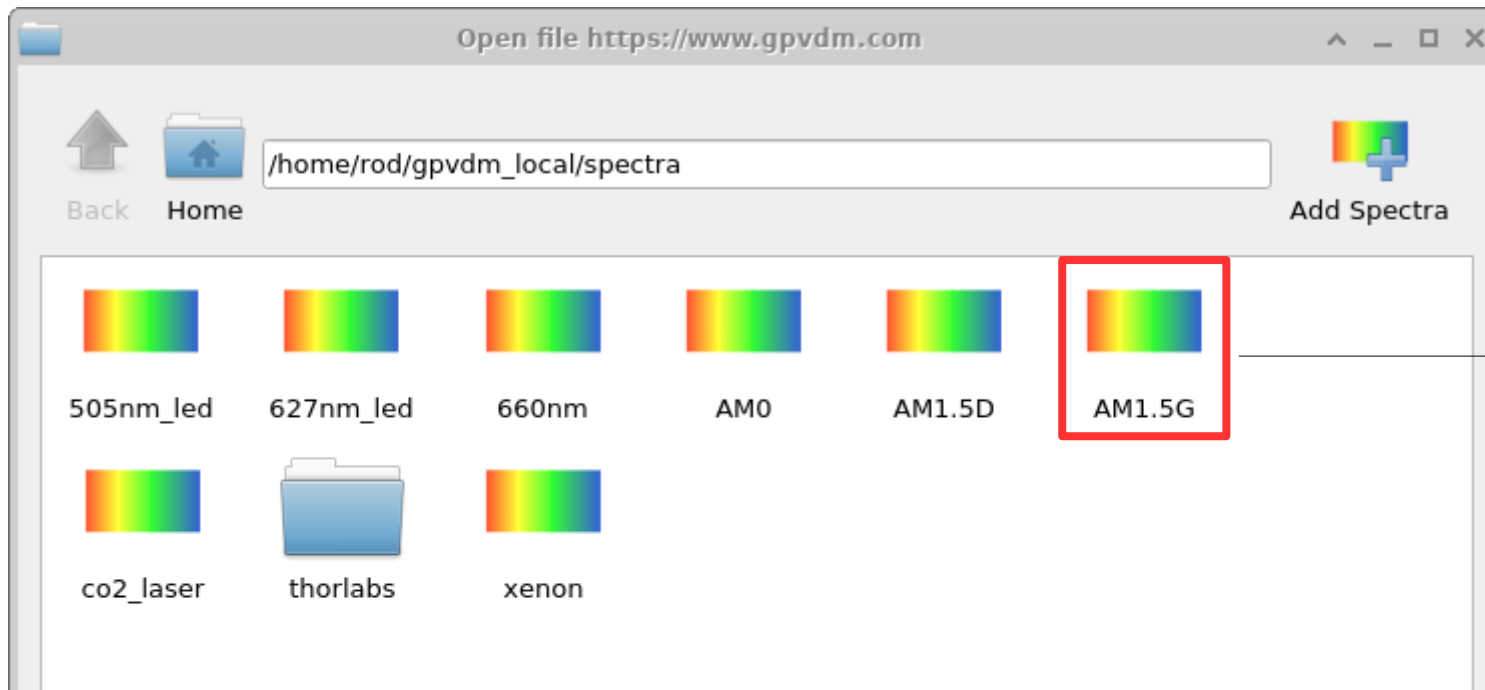
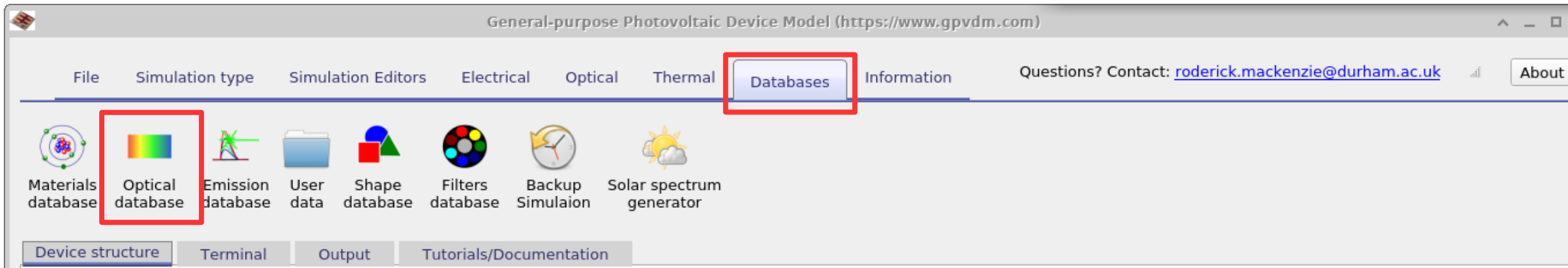
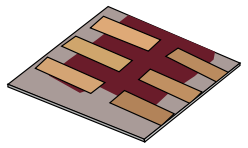
- White LEDs



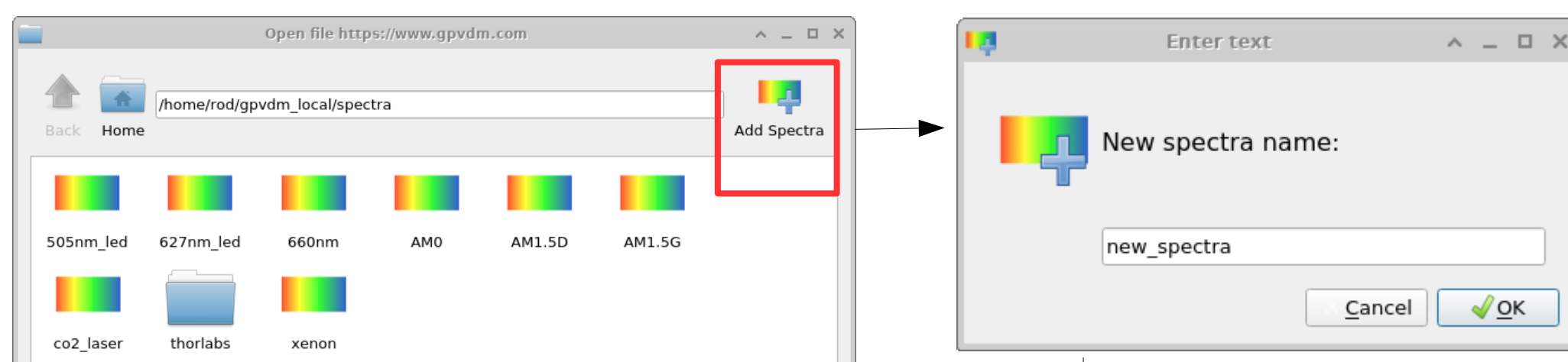
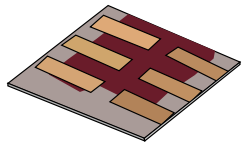
[https://commons.wikimedia.org/wiki/File:Red-YellowGreen-Blue\\_LED\\_spectra.png](https://commons.wikimedia.org/wiki/File:Red-YellowGreen-Blue_LED_spectra.png)



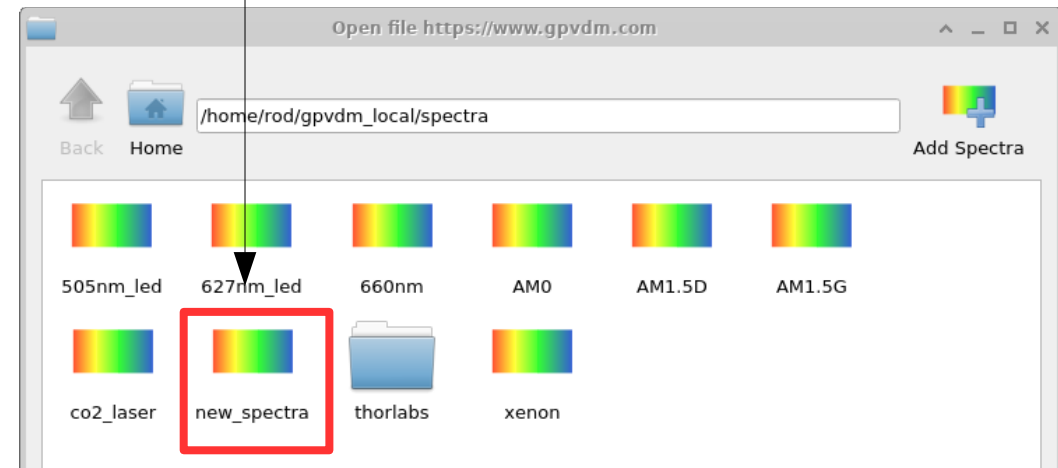
# You can access the spectra in from the database tab:



# Exercise 1: Sometimes you will want to import your own spectra say from a lamp or other source in your lab

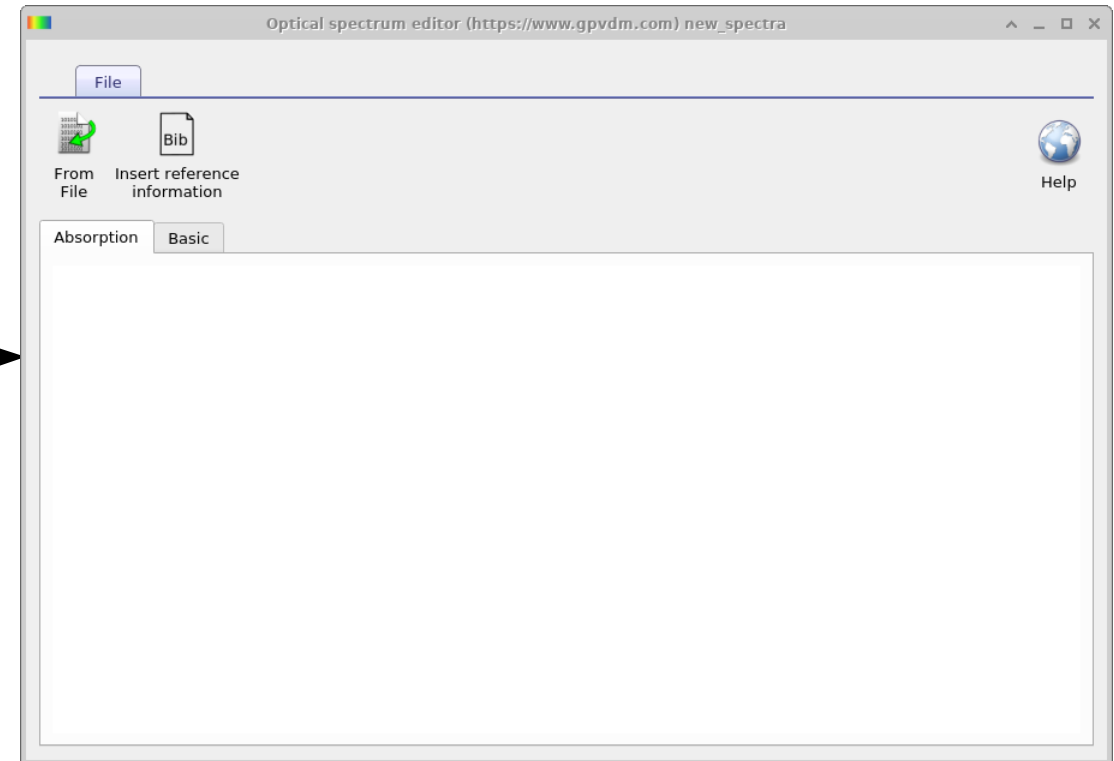
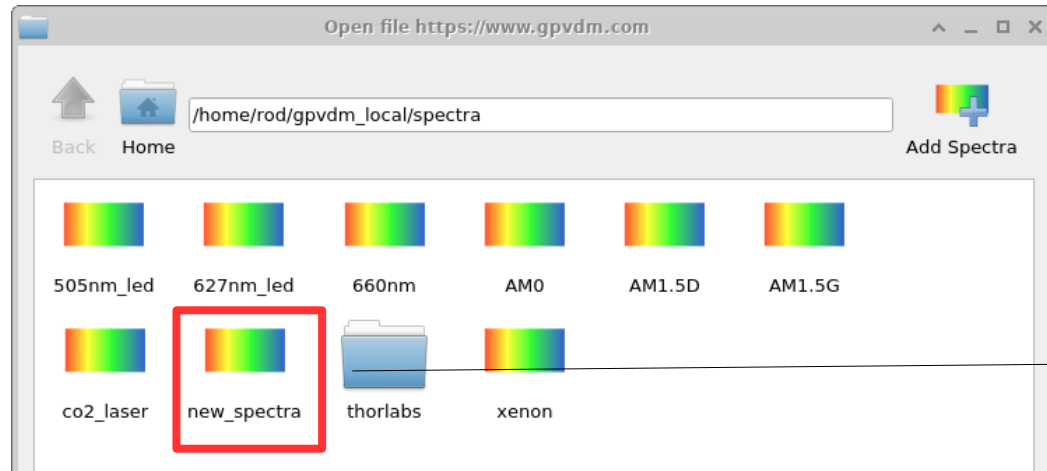
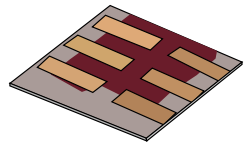


- Click on the “Add spectra” button
- Type a new name
- Open the new spectra by double clicking.



# Exercise 1:

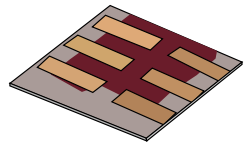
## Opening your new spectrum window



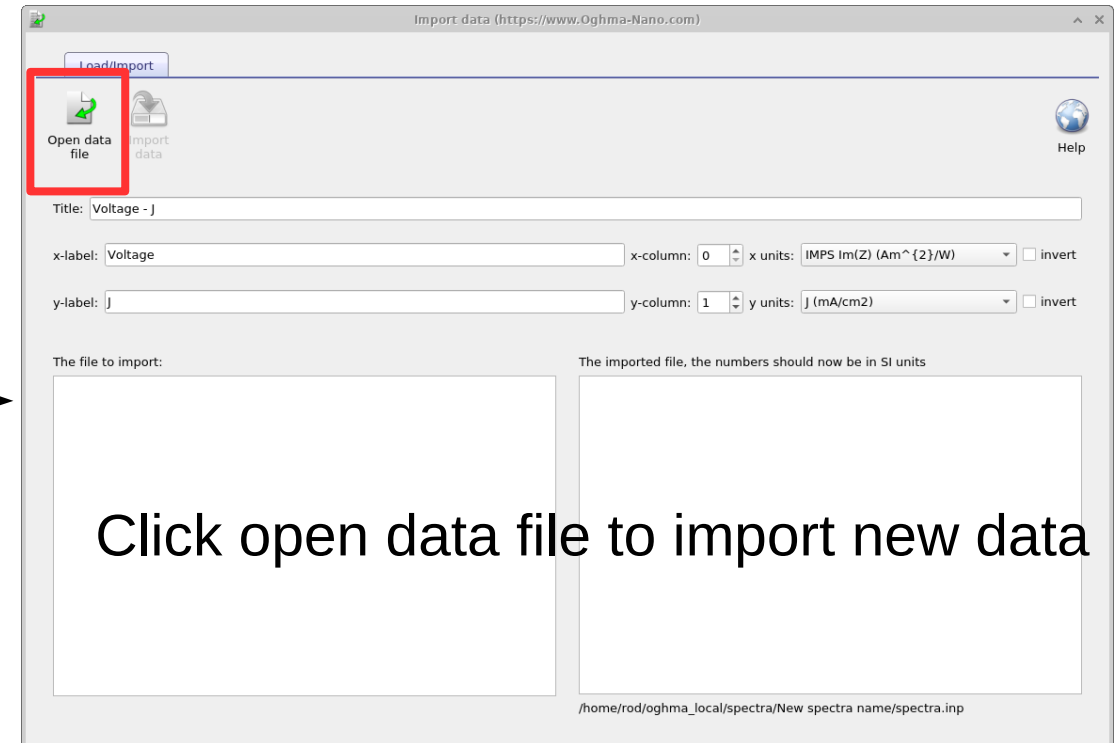
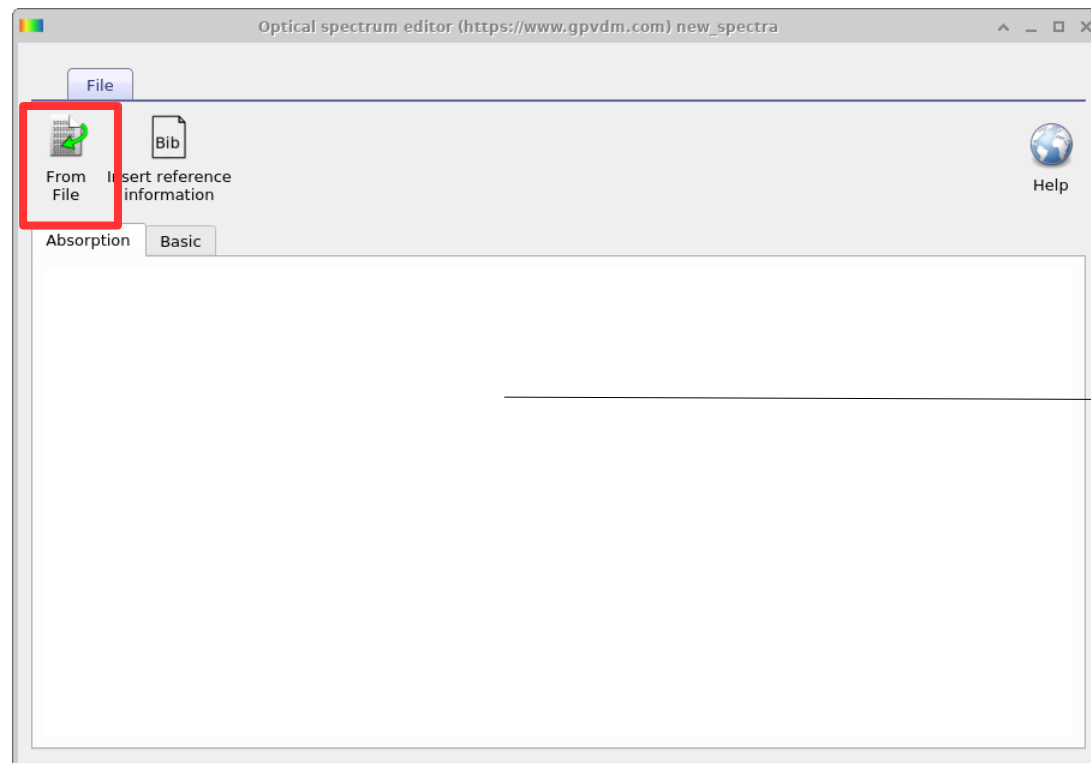
- This will open an empty spectrum window.
- We now need some data to fill it.

# Exercise 1:

## Importing a spectra



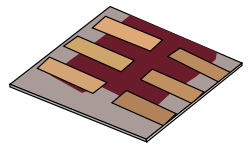
- Download the solar spectra of the sun from:
- [https://www.oghma-nano.com/demo/solar\\_spectra\\_demo.txt](https://www.oghma-nano.com/demo/solar_spectra_demo.txt)
- And save it to your home directory.



- Then open it using “File Import” (red box below)

# Exercise 1:

## Your imported data will appear like this:



Import data (https://www.Oghma-Nano.com)

Load/Import

Open data file Import data Help

Title: Wavelength - Intensity

x-label: Wavelength x-column: 0 x units: Wavelength (nm) ☐ invert

y-label: Intensity y-column: 1 y units: Intensity (nm<sup>-1</sup>.Wm<sup>-2</sup>) ☐ invert

The file to import:

#Wavelength nm	Direct+circumsolar W*m-2*nm-1
280	2.5361E-26
280.5	1.0917E-24
281	6.1253E-24
281.5	2.7479E-22
282	2.8346E-21
282.5	1.3271E-20
283	6.7646E-20
283.5	1.4614E-19
284	4.9838E-18
284.5	2.1624E-17
285	8.9998E-17
285.5	6.4424E-16
286	2.3503E-15
286.5	1.8458E-14
287	7.2547E-14
287.5	3.6618E-13
288	2.8061E-12
288.5	9.0651E-12
289	3.4978E-11

/home/rod/Desktop/sun.txt

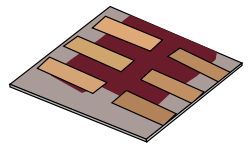
The imported file, the numbers should now be in SI units

#oghma_csv {"title":"Wavelength - Intensity","type":"xy","y_label":"Wavelength"	
2.800000e-07	2.536100e-17
2.805000e-07	1.091700e-15
2.810000e-07	6.125300e-15
2.815000e-07	2.747900e-13
2.820000e-07	2.834600e-12
2.825000e-07	1.327100e-11
2.830000e-07	6.764600e-11
2.835000e-07	1.461400e-10
2.840000e-07	4.983800e-09
2.845000e-07	2.162400e-08
2.850000e-07	8.999800e-08
2.855000e-07	6.442400e-07
2.860000e-07	2.350300e-06
2.865000e-07	1.845800e-05
2.870000e-07	7.254700e-05
2.875000e-07	3.661800e-04
2.880000e-07	2.806100e-03
2.885000e-07	9.065100e-03

/home/rod/oghma\_local/spectra/new\_spectra/spectra.inp

# Exercise 1:

## Closer look at the data importer tool



Import data (https://www.Oghma-Nano.com)

Load/Import

Open data file Import data Help

Title: Wavelength - Intensity

x-label: Wavelength x-column: 0 x units: Wavelength (nm) ☐ invert

y-label: Intensity y-column: 1 y units: Intensity (nm<sup>-1</sup>.Wm<sup>-2</sup>) ☐ invert

The file to import:

#	Wavelength nm	Direct+circumsolar W*m-2*nm-1
280	2.5361E-26	
280.5	1.0917E-24	
281	6.1253E-24	
281.5	2.7479E-22	
282	2.8346E-21	
282.5	1.3271E-20	
283	6.7646E-20	
283.5	1.4614E-19	
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284.5	2.1624E-17	
285	8.9998E-17	
285.5	6.4424E-16	
286	2.3503E-15	
286.5	1.8458E-14	
287	7.2547E-14	
287.5	3.6618E-13	
288	2.8061E-12	
288.5	9.0651E-12	
289	3.4978E-11	

Raw data from FILE not in SI

The imported file: The numbers should now be in SI units

#	oghma	Wavelength (nm)	Intensity (nm <sup>-1</sup> .Wm <sup>-2</sup> )
280	2.536100e-26	2.536100e-17	
280.5	1.091700e-24	1.091700e-15	
281	6.125300e-24	6.125300e-15	
281.5	2.747900e-22	2.747900e-13	
282	2.834600e-21	2.834600e-12	
282.5	1.327100e-20	1.327100e-11	
283	6.764600e-20	6.764600e-11	
283.5	1.461400e-19	1.461400e-10	
284	4.983800e-18	4.983800e-09	
284.5	2.162400e-17	2.162400e-08	
285	8.999800e-17	8.999800e-08	
285.5	6.442400e-16	6.442400e-07	
286	2.350300e-15	2.350300e-06	
286.5	1.845800e-14	1.845800e-05	

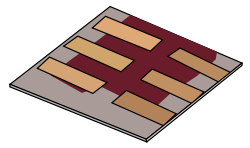
Converted to SI.

x-column: 0 x units: Wavelength (nm) ☐ invert

y-column: 1 y units: Intensity (nm<sup>-1</sup>.Wm<sup>-2</sup>) ☐ invert

/home/rod/Desktop/sun.tx

- The model needs all data to be in SI units.
- The data you downloaded had units of Wavelength (nm) v.s. Intensity (nm<sup>-1</sup>Wm<sup>-2</sup>). Intensity is already SI, wavelength in nm is not.
- To convert the data to SI, select from the drop down boxes what units the INPUT DATA is in.
- The converted data will be displayed in the right hand text box.



# Exercise 1:

## Closer look at the data importer tool

Import data (https://www.Oghma-Nano.com)

Load/Import

Open data file Import data Help

Title: Wavelength - Intensity

x-label: Wavelength x-column: 0 x units: Wavelength (nm) ☐ invert

y-label: Intensity y-column: 1 y units: Intensity (nm<sup>-1</sup>.Wm<sup>-2</sup>) ☐ invert

The file to import:

#	Wavelength nm	Direct+circumsolar W*m-2*nm-1
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283	6.7646E-20	
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285	8.9998E-17	
285.5	6.4424E-16	
286	2.3503E-15	
286.5	1.8458E-14	
287	7.2547E-14	
287.5	3.6618E-13	
288	2.8061E-12	
288.5	9.0651E-12	
289	3.4978E-11	

Raw data from FILE not in SI

The imported file: the numbers should now be in SI units

#	oghma	Wavelength - Intensity
280	2.536100e-17	
280.5	1.091700e-15	
281	6.125300e-15	
281.5	2.747900e-13	
282	2.834600e-12	
282.5	1.327100e-11	
283	6.764600e-11	
283.5	1.461400e-10	
284	4.983800e-09	
284.5	2.162400e-08	
285	8.999800e-08	
285.5	6.442400e-07	
286	2.350300e-06	
286.5	1.845800e-05	

Converted to SI.

x-column: 0 x units: Wavelength (nm) ☐ invert

y-column: 1 y units: Intensity (nm<sup>-1</sup>.Wm<sup>-2</sup>) ☐ invert

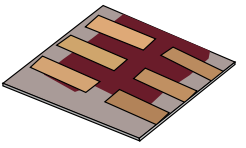
/home/rod/Desktop/sun.tx

• You should see on the right hand side of the window nm has been converted into units of meters.

• Always perform a sanity check on the imported data by looking at the numbers in each column – do they make sense?

• They should always be in SI units. There should be no Inf values, and no NaN values if anything

# Try at home demo



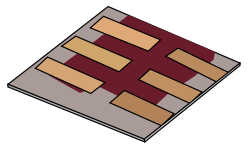
- You can download the data used in this example from:

[https://www.oghma-nano.com/demo/solar\\_spectra\\_demo.txt](https://www.oghma-nano.com/demo/solar_spectra_demo.txt)

- Have a go at importing the data your self.

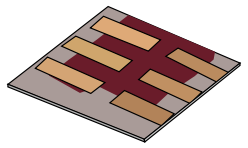


# Outline of the talk

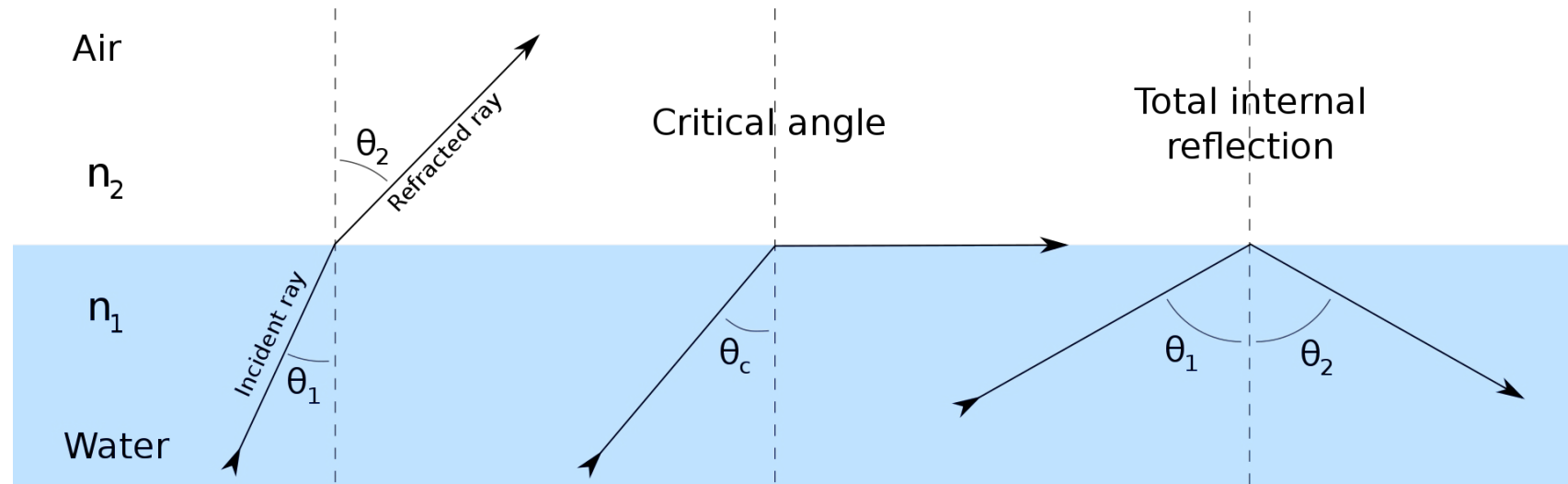
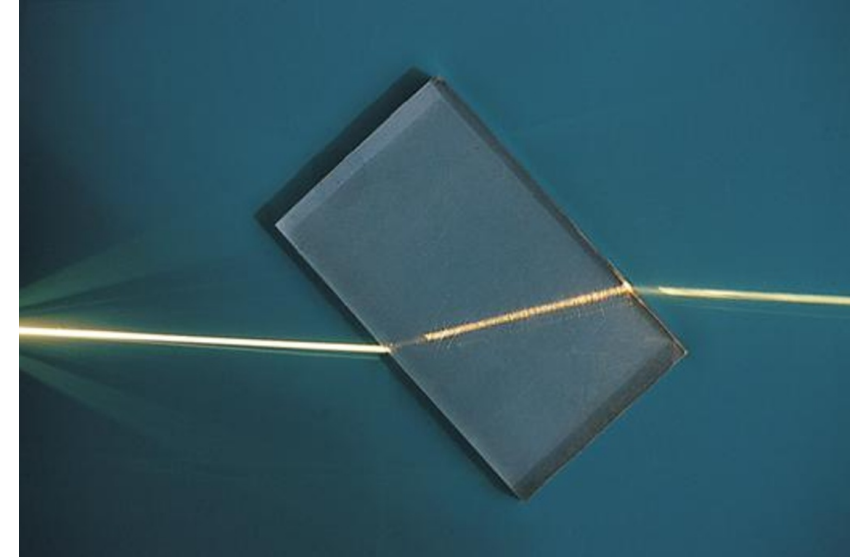


- In this talk we will cover:
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  - **What you need for accurate optical simulations**
    - » Optical spectra
    - » **Refractive index data (n)**
    - » Optical absorption data (k)
      - The materials database
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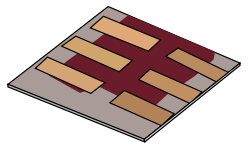
# Refractive index: Snells law and the speed of light



- Refractive index governs by how much light is bent when entering or leaving an object.
- Snell's law describes this bending of light:
$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$
- Refractive index also governs the speed of light in an object.



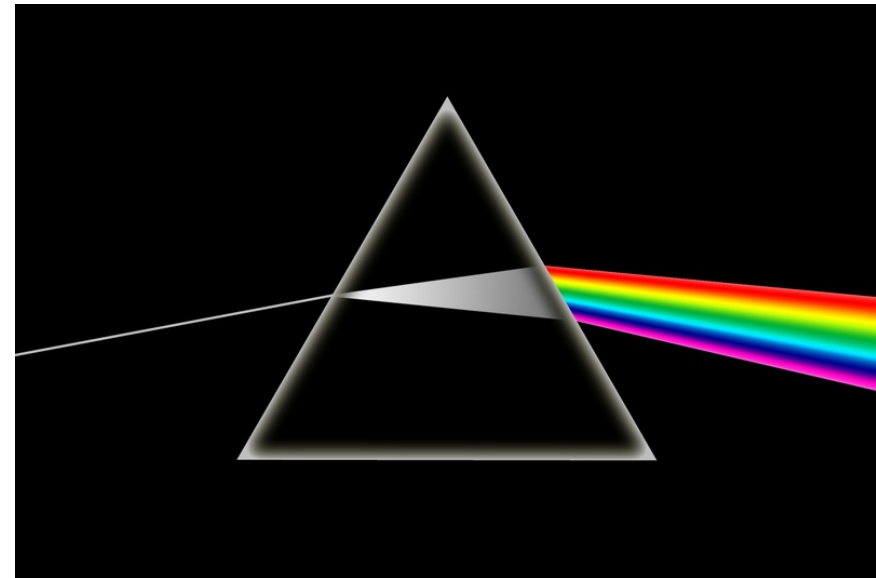
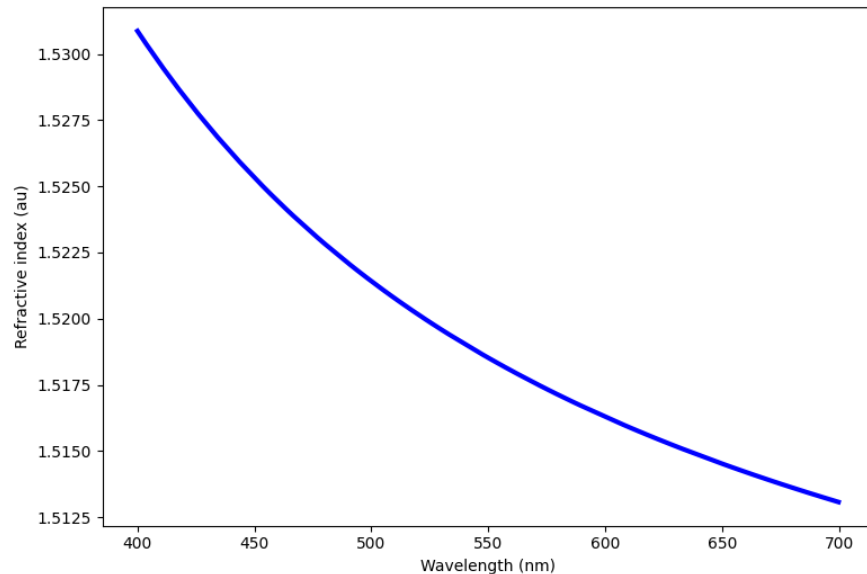
# Refractive index: Variation as a function of wavelength



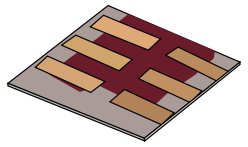
- The reason you see light being split in a prism is due to a combination of Snells law and materials having different refractive index values at different wavelengths.

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

- Below is a plot of the refractive index of glass as a function of wavelength.



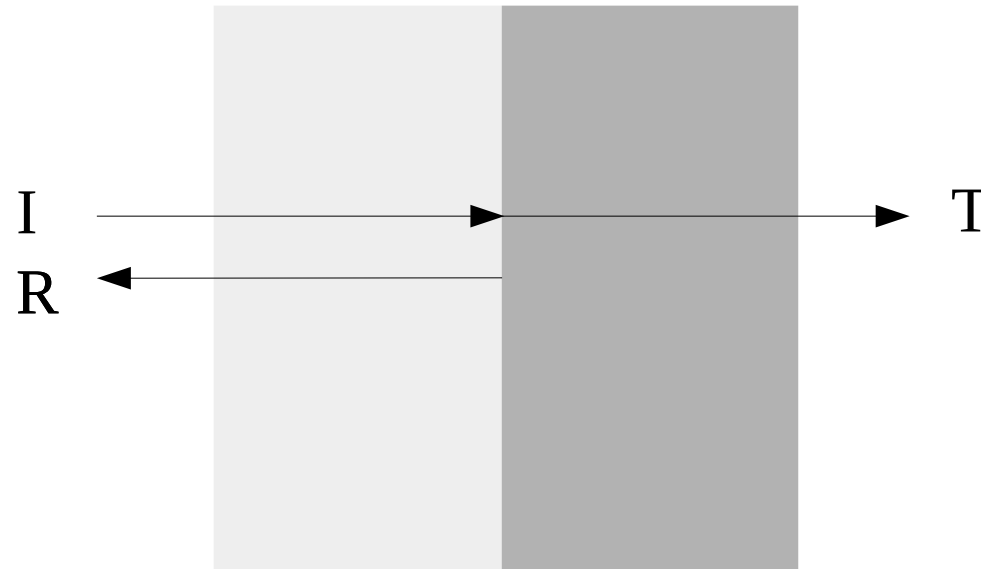
# Refractive index (n): Reflection



- Refractive index also governs how much light is reflected/transmitted at an interface:

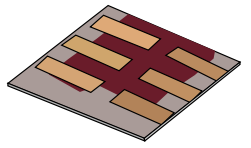
$$R = \left| \frac{n_1 - n_2}{n_1 + n_2} \right|^2$$

$$T = 1 - R$$



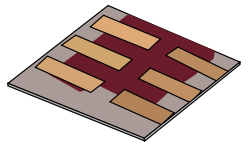
- For this talk I'm not going to go further into Snell's law or these relations, I just want you to appreciate the how refractive index influences optical simulations.

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# Optical absorption

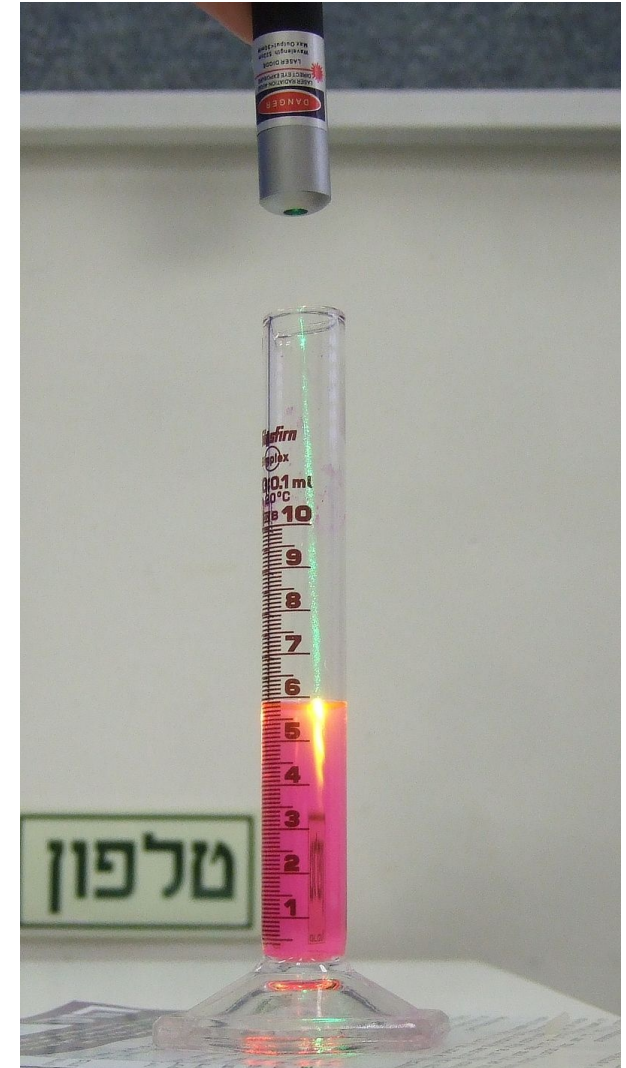
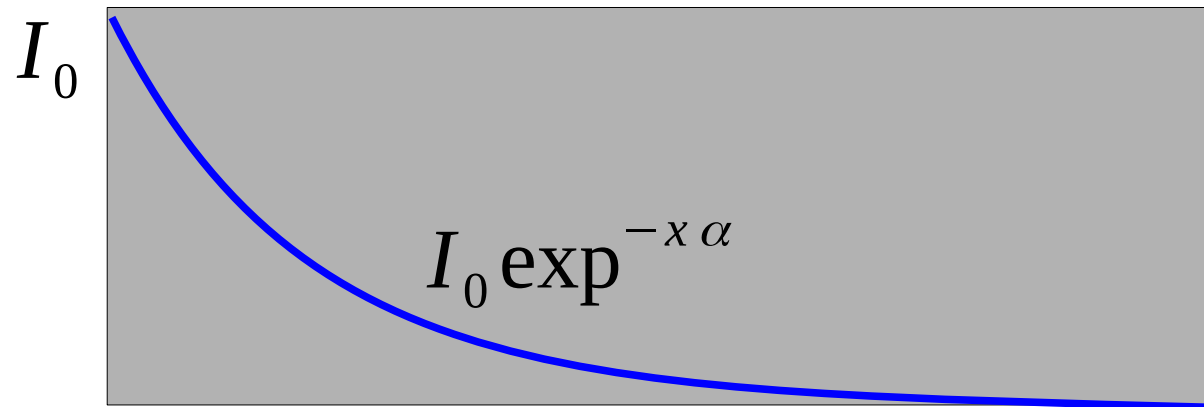


- Optical absorption is the process where by light is absorbed as it passes through a material.

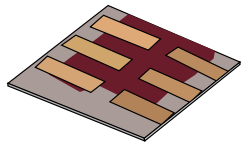
- If follows the equation

$$I(x) = I_0 \exp^{-x \alpha}$$

- Where  $I_0$  is the initial photon flux,  $I(x)$  is the photon flux at position  $x$  in the medium and  $\alpha$  is the absorption coefficient.



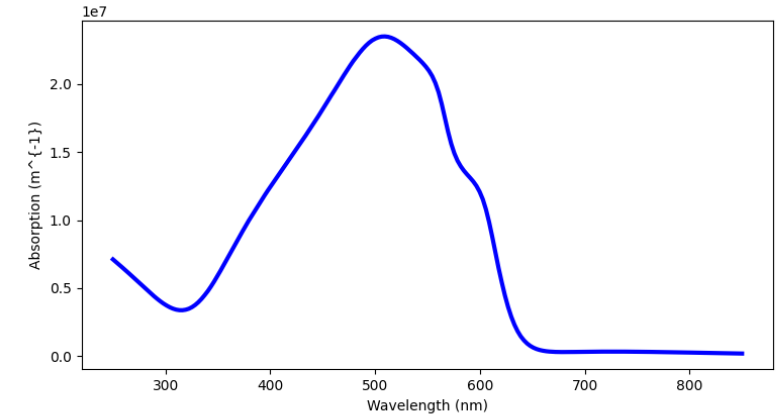
# Optical absorption: Absorption coefficient (alpha)



- The alpha in this equation also changes as a function of wavelength, an example is given to the left for PTB7 a commonly used polymer in organic electronics.

- So the equation

$$I(x) = I_0 \exp^{-x \alpha}$$

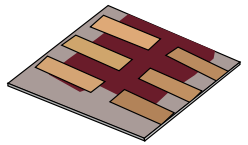


- Should really be written as a function of wavelength:

$$I(x, \lambda) = I_0 \exp^{-x \alpha(\lambda)}$$

- Alpha has units of length<sup>-1</sup>, in this case as the model works only in SI, m<sup>-1</sup>.

# Optical absorption: Attenuation coefficient (k)



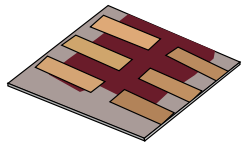
- As a final note you often hear about **n/k** data
- This is a complex number representing both the n term and the alpha term:
- This is written as  $\bar{n} = n + j \kappa = n + j \frac{\lambda \alpha}{4 \pi}$
- The relationship between alpha and k is given as  $\alpha = \frac{4 \pi \kappa}{\lambda}$
- The model takes **alpha** (m<sup>-1</sup>) as an input, so using the above formula you can convert values of k found in the literature for use in the model.
- The model also has an option to convert k to alpha in the import window:

x-column: 0 x units: Wavelength (nm) ☐ invert

y-column: 2 y units: **Attenuation coefficient (au)** ☐ invert



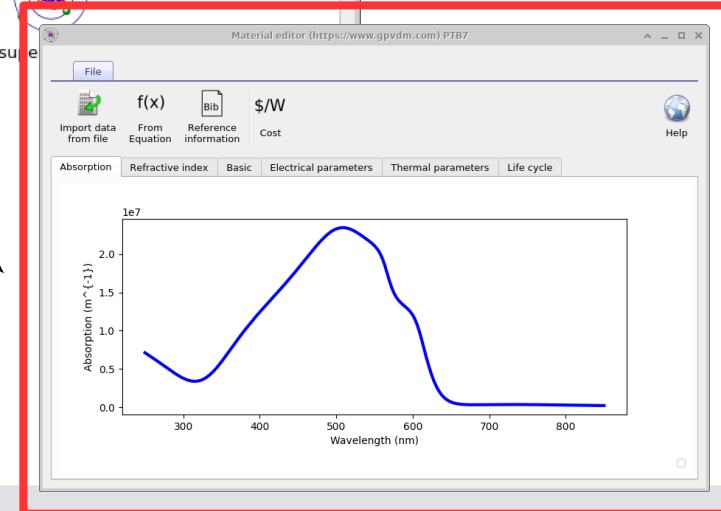
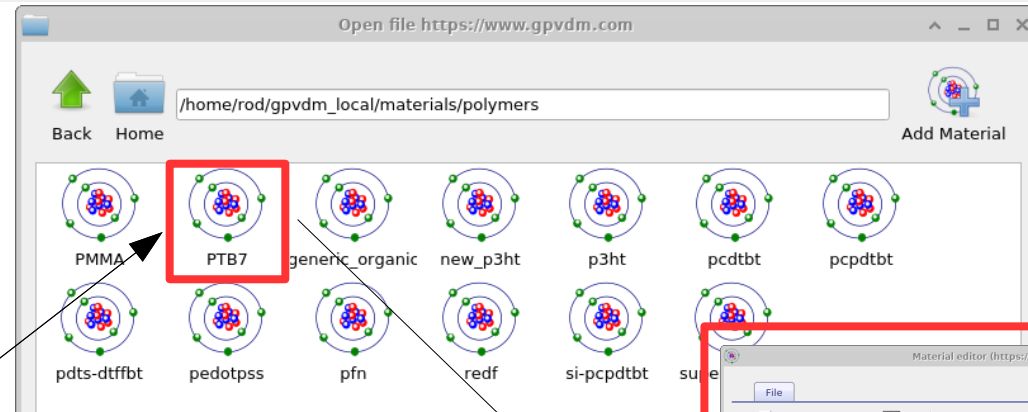
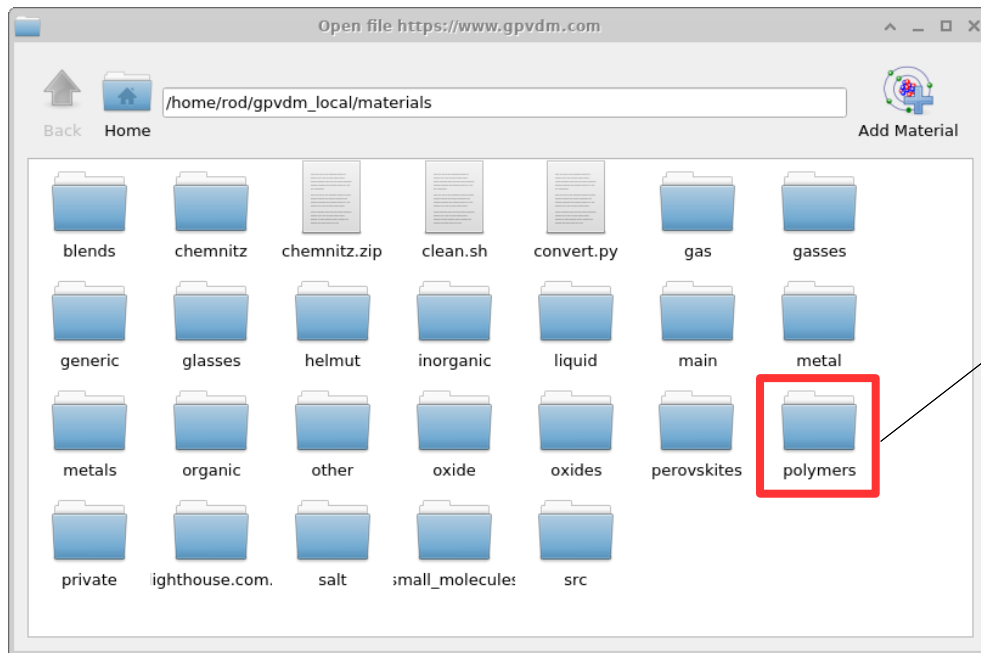
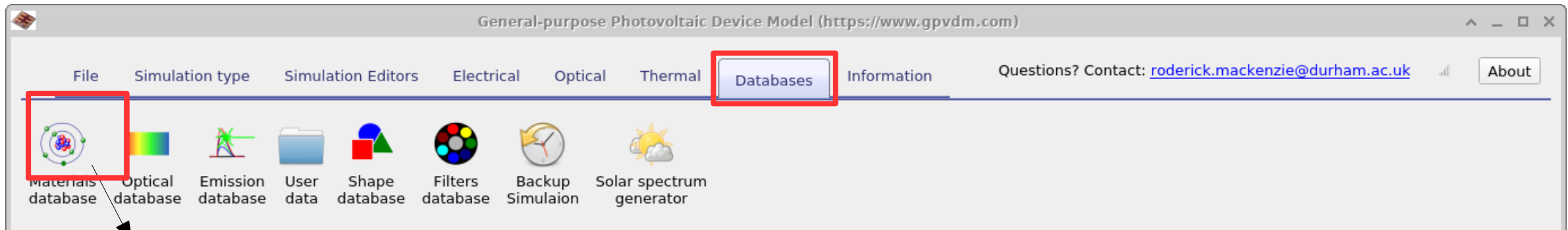
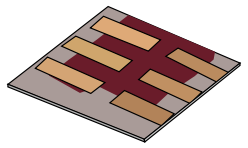
# Outline of the talk



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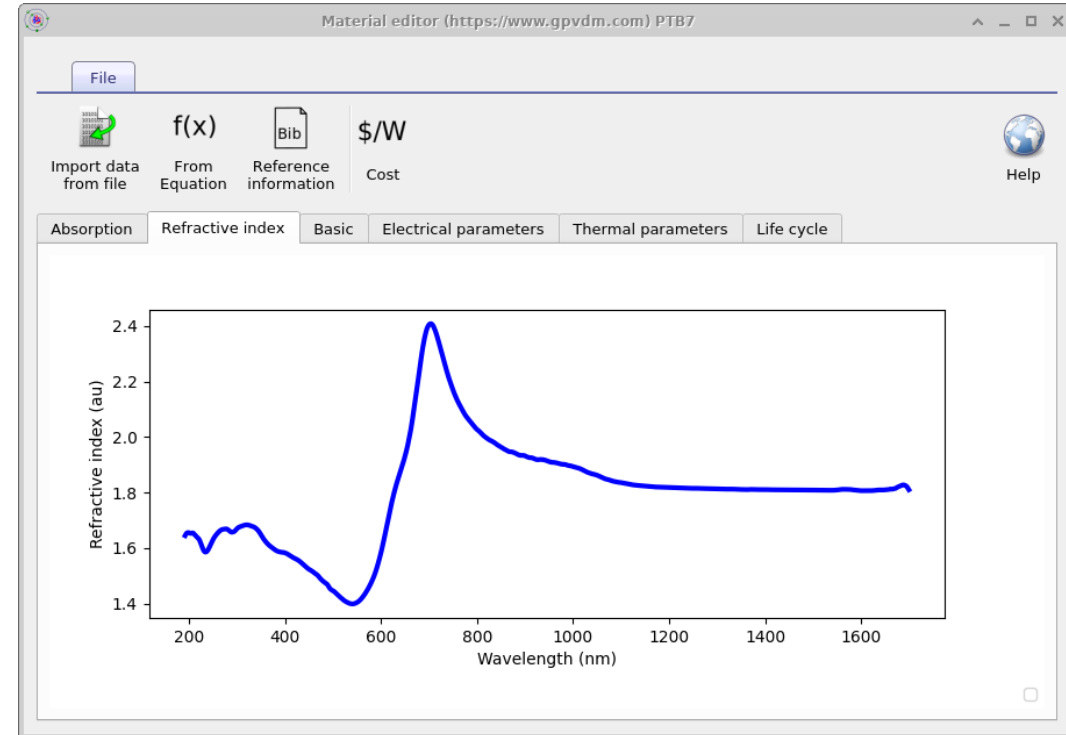
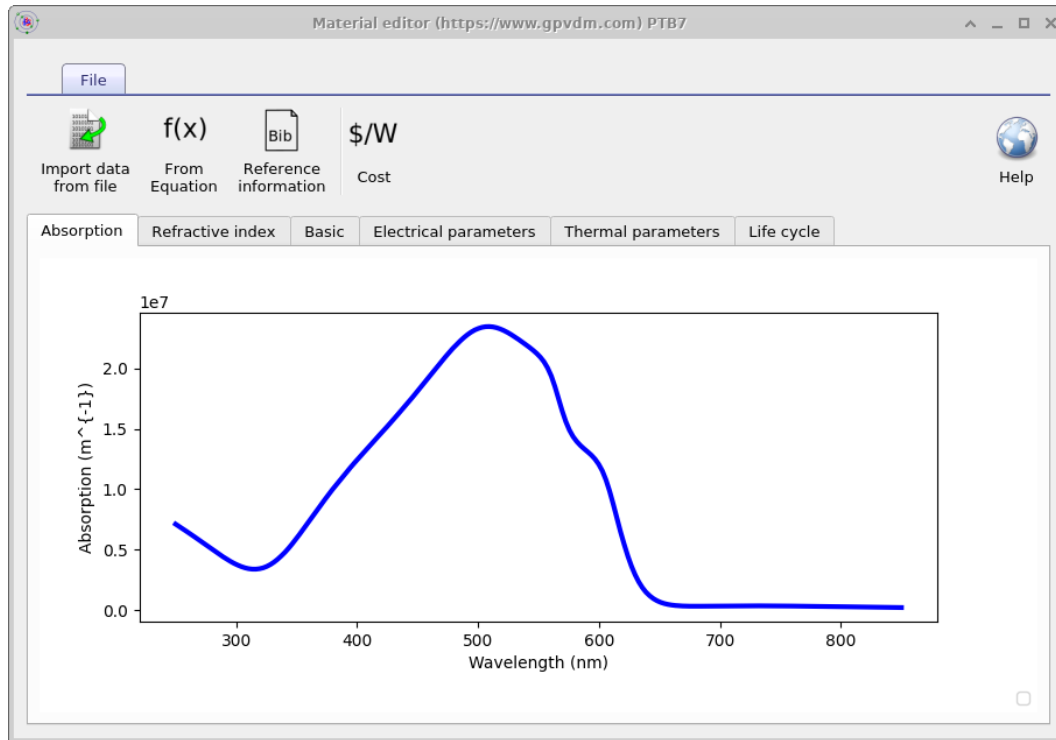
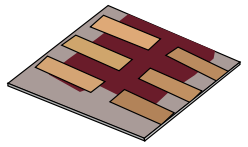
# The materials database:

## This has lots of materials in it that you can explore



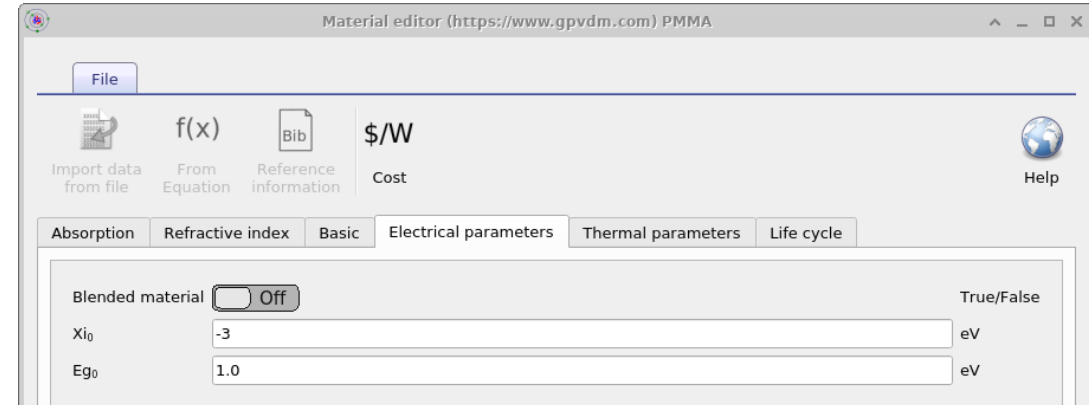
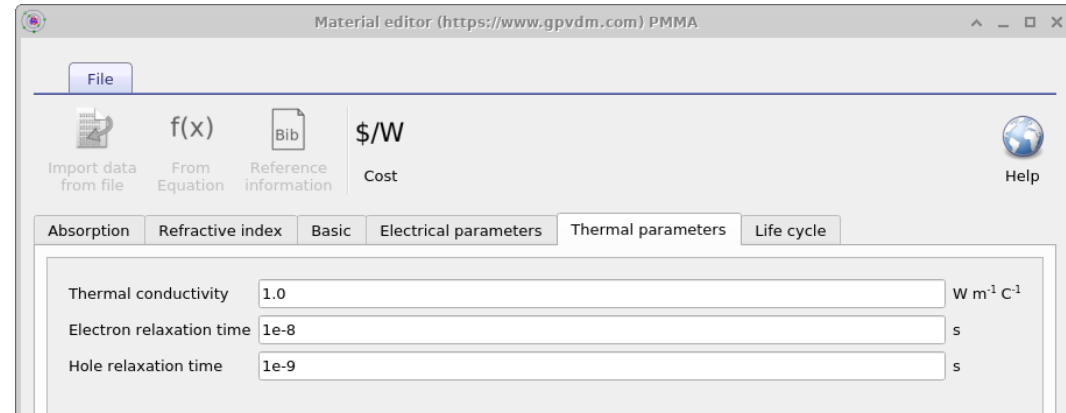
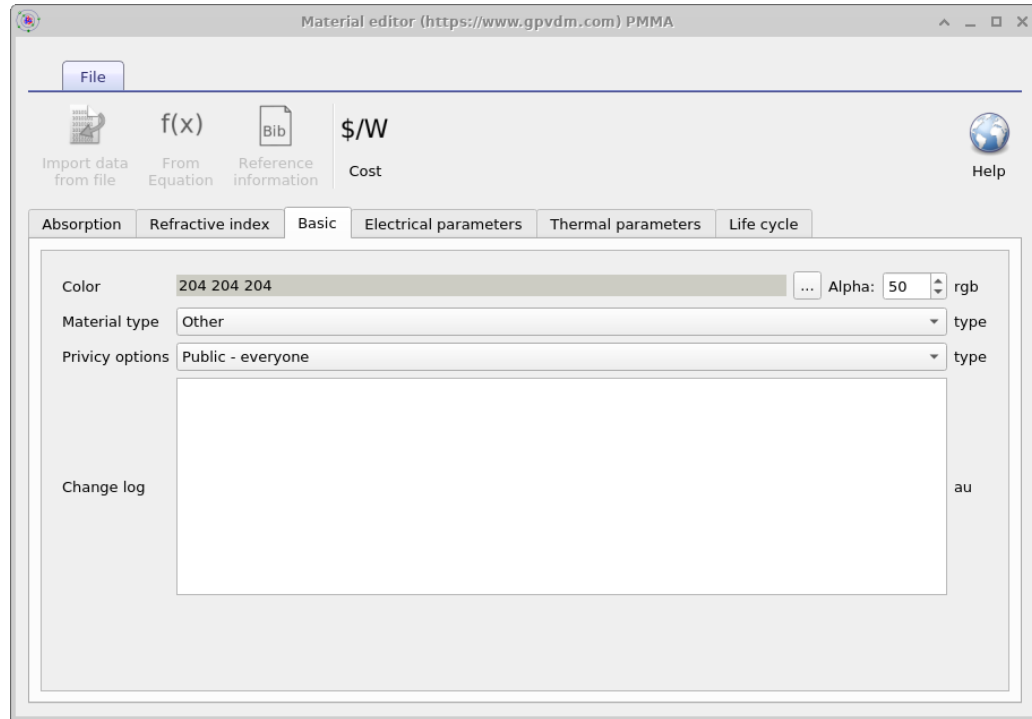
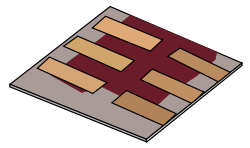
- Each material has  $n/\alpha$  data stored for it.

# The materials database: n/alpha



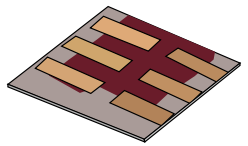
- Each material has both refractive index and absorption data associated with it.

# The materials database: Other parameters



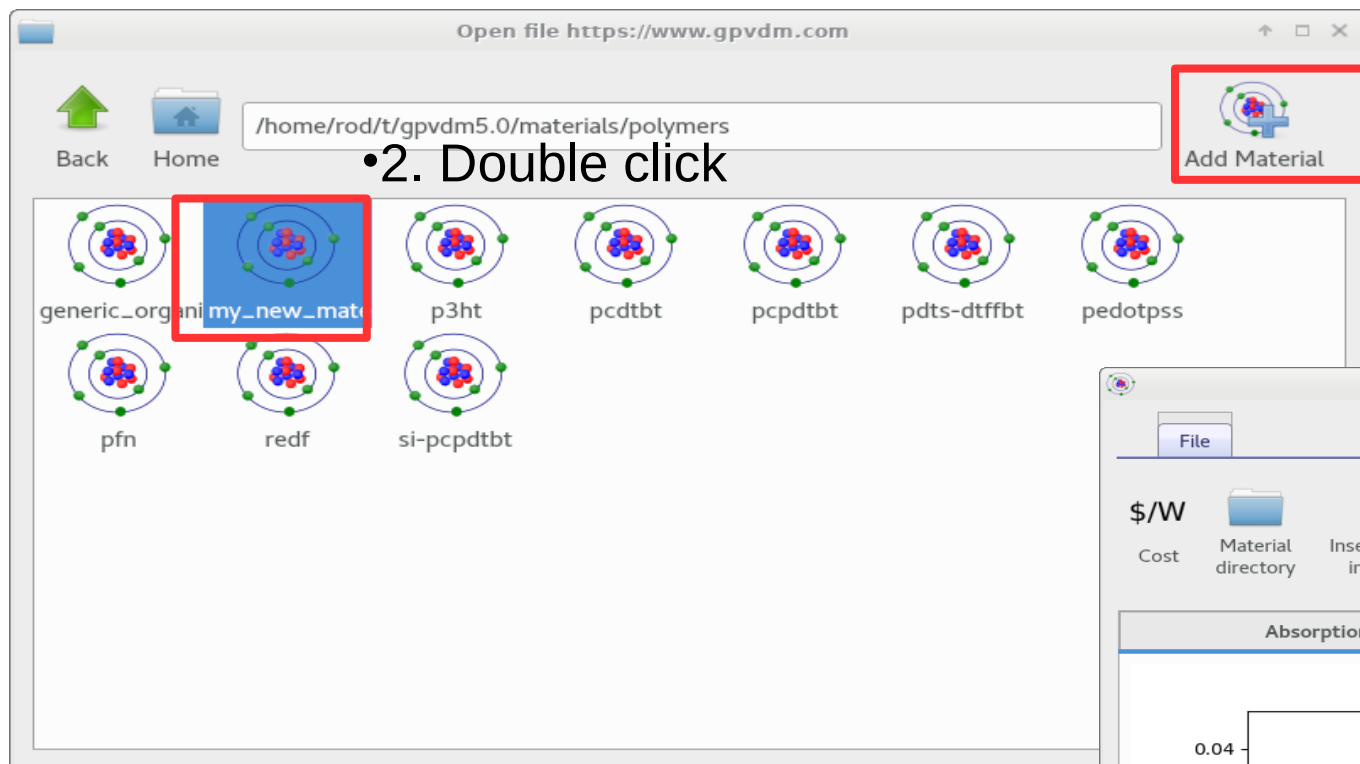
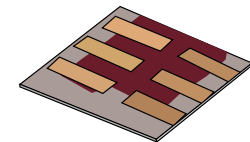
- Other tabs show other basic material parameters

# Outline of the talk



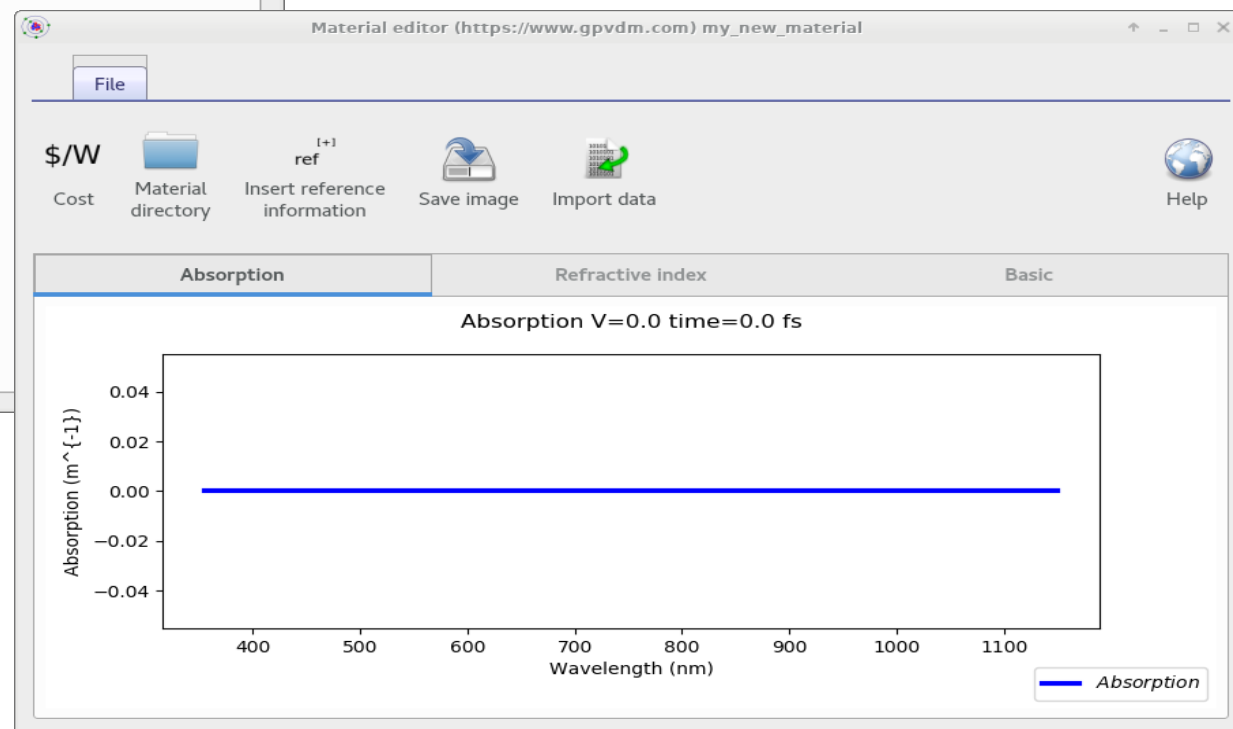
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# Making a new material



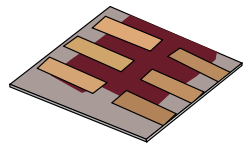
•1. Click

•2. Double click



•The material file has no data....

# Let's get hold of some data to import... usually you would get this from experiment or a publication.



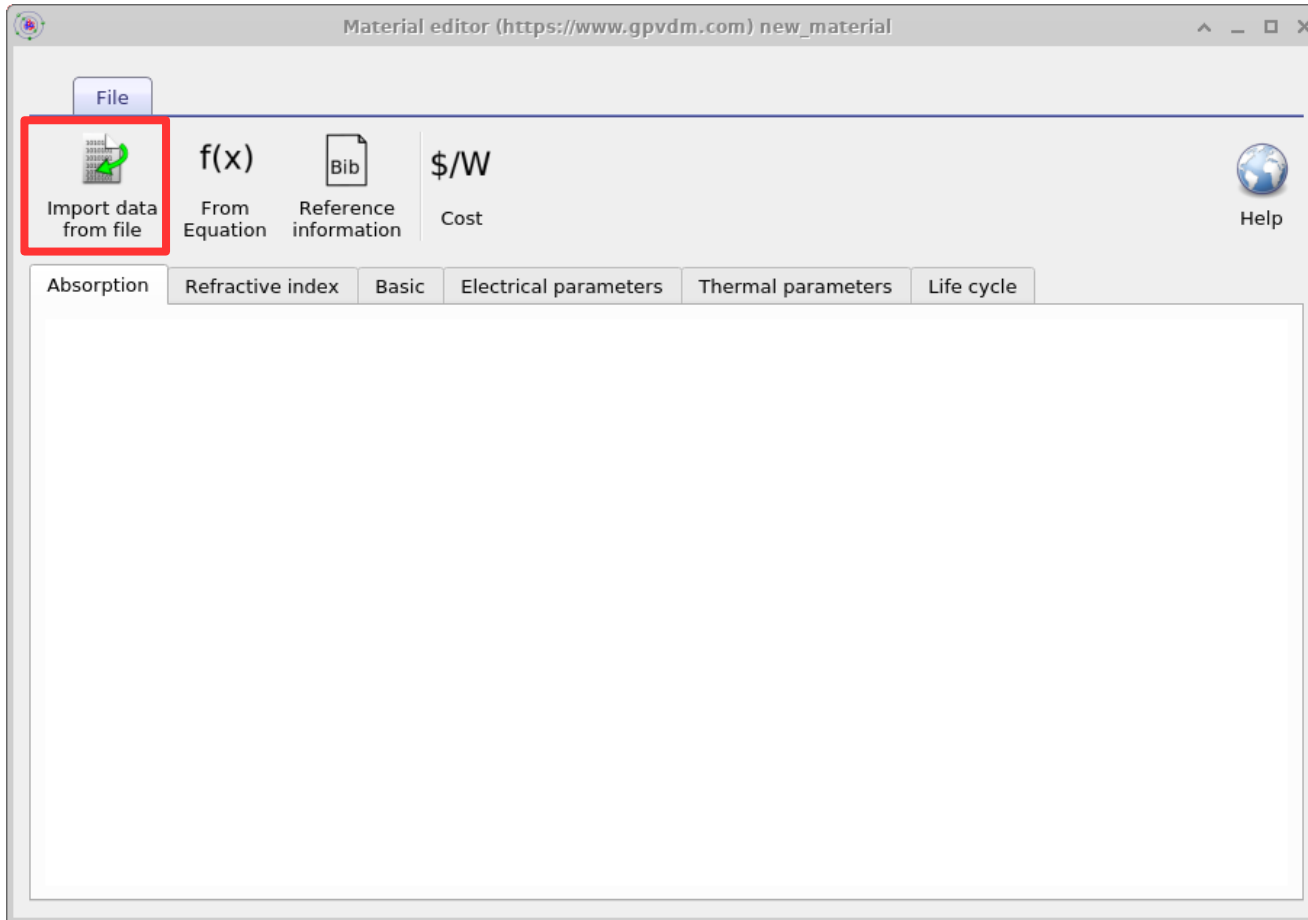
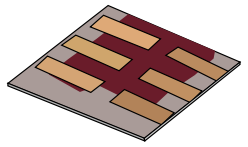
- Download this zip archive:

[https://www.oghma-nano.com/demo/ptb7pc70bm\\_demo.zip](https://www.oghma-nano.com/demo/ptb7pc70bm_demo.zip)

```
1#wavelength (nm), refractive index
2 305.79 1.9154
3 316.66 1.90433
4 323.66 1.88767
5 329.09 1.87657
6 335.31 1.85712
7 336.86 1.85713
8 346.96 1.83492
9 356.29 1.81549
10 363.28 1.8044
11 364.06 1.79049
12 371.04 1.80445
13 377.23 1.81841
14 395.86 1.81574
15 398.18 1.81593
16 405.16 1.82415
17 421.48 1.78806
18 422.25 1.78613
19 426.13 1.7853
```

```
1#wavelength (nm), k (complex refractive index)
2 301.31 0.38976
3 305.2 0.37308
4 308.31 0.35361
5 309.88 0.33692
6 313.77 0.32024
7 315.32 0.3119
8 321.55 0.28688
9 326.99 0.273
10 329.32 0.26744
11 341.74 0.25082
12 347.95 0.24807
13 354.16 0.24811
14 358.81 0.25092
15 362.69 0.25373
16 367.34 0.26211
17 373.54 0.2705
18 377.42 0.27052
19 382.85 0.26777
```

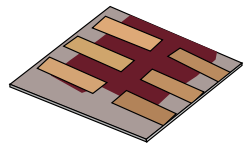
# Let's import k.csv



- Select “Import the data from file” button.
- Select k.csv to import... from where ever you extracted n/k.csv
- Again the import window will pop up.



The raw data will be on the LHS of the import window  
The RHS will display the data converted to SI units.



Import data (https://www.Oghma-Nano.com)

Load/Import

Open data file Import data Help

Title: Wavelength - Absorption

x-label: Wavelength x-column: 0 x units: Wavelength (nm) invert

y-label: Absorption y-column: 1 y units: Attenuation coefficient (au) invert

The file to import:

```
#wavelength (nm), k (complex refractive index)
301.31 0.38976
305.2 0.37308
308.31 0.35361
309.88 0.33692
313.77 0.32024
315.32 0.3119
321.55 0.28688
326.99 0.273
329.32 0.26744
341.74 0.25082
347.95 0.24807
354.16 0.24811
358.81 0.25092
362.69 0.25373
367.34 0.26211
373.54 0.2705
377.42 0.27052
382.85 0.26777
393.72 0.26227
```

•The imported data

The imported file, the numbers should now be in SI units

```
#oghma_csv {"title":"Wavelength - Absorption","type":"xy","y_label":"Wavelength - Absorption"}
3.013100e-07 1.625523e+07
3.052000e-07 1.536126e+07
3.083100e-07 1.441212e+07
3.098800e-07 1.366290e+07
3.137700e-07 1.282548e+07
3.153200e-07 1.221000e+07
3.215500e-07 1.121144e+07
3.269900e-07 1.049150e+07
3.293200e-07 1.020810e+07
3.417400e-07 9.223077e+06
3.479500e-07 8.99152e+06
3.541600e-07 8.81507e+06
3.588100e-07 8.787801e+06
3.626900e-07 8.791151e+06
3.673400e-07 8.966539e+06
3.735400e-07 9.099964e+06
3.774200e-07 9.007079e+06
3.828500e-07 8.789067e+06
```

•The the data transformed into SI, ready to be imported into the model.

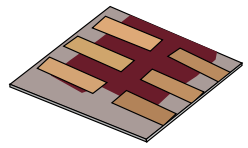
•Set these to the units used in the RAW data file you imported.

x-column: 0 x units: Wavelength (nm) invert

y-column: 1 y units: Attenuation coefficient (au) invert

•Set the values to **Wavelength (nm)**, and **Attenuation coefficient (au)**, the RHS will then be in SI, scroll down to inspect the file. Make sure it has been imported correctly

# Then click import data and the data will be imported into the material..



Import data (https://www.Oghma-Nano.com)

Load/Import

Open data file Import data Help

Title: Wavelength - Absorption

x-label: Wavelength x-column: 0 x units: Wavelength (nm) ☐ invert

y-label: Absorption y-column: 1 y units: Attenuation coefficient (au) ☐ invert

The file to import:

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301.31 0.38976
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347.95 0.24807
354.16 0.24811
358.81 0.25092
362.69 0.25373
367.34 0.26211
373.54 0.2705
377.42 0.27052
382.85 0.26777
393.72 0.26227

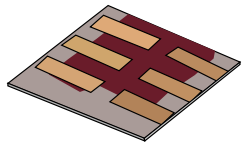
/home/rod/Desktop/k.csv

The imported file, the numbers should now be in SI units

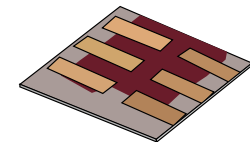
#oghma_csv {"title": "Wavelength - Absorption", "type": "xy", "y_label": "Wavelength - Absorption", "x_label": "Wavelength (nm)", "y_units": "Attenuation coefficient (au)", "x_units": "Wavelength (nm)"}	
3.013100e-07	1.625523e+07
3.052000e-07	1.536126e+07
3.083100e-07	1.441274e+07
3.098800e-07	1.366290e+07
3.137700e-07	1.282548e+07
3.153200e-07	1.243006e+07
3.215500e-07	1.121144e+07
3.269900e-07	1.049150e+07
3.293200e-07	1.020511e+07
3.417400e-07	9.223077e+06
3.479500e-07	8.959152e+06
3.541600e-07	8.803477e+06
3.588100e-07	8.787801e+06
3.626900e-07	8.791151e+06
3.673400e-07	8.966539e+06
3.735400e-07	9.099964e+06
3.774200e-07	9.007079e+06
3.828500e-07	8.789067e+06

/home/rod/oghma\_local/materials/fred/alpha.csv

# Outline of the talk

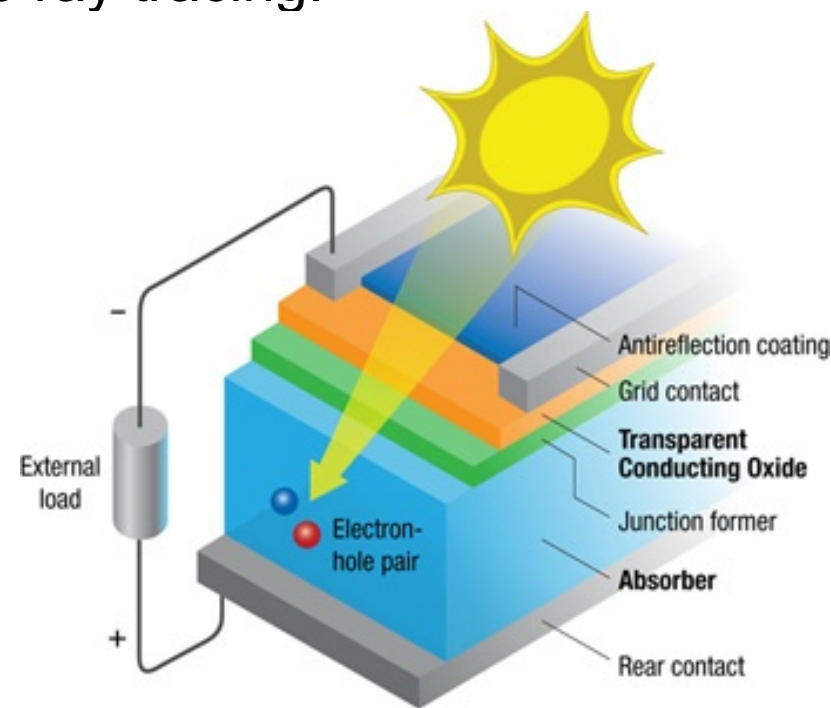


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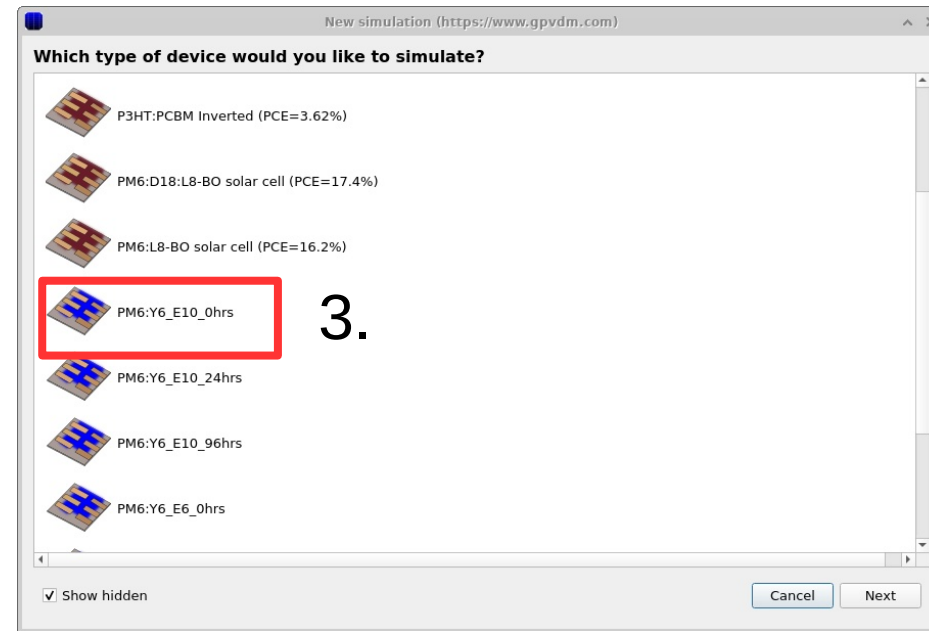
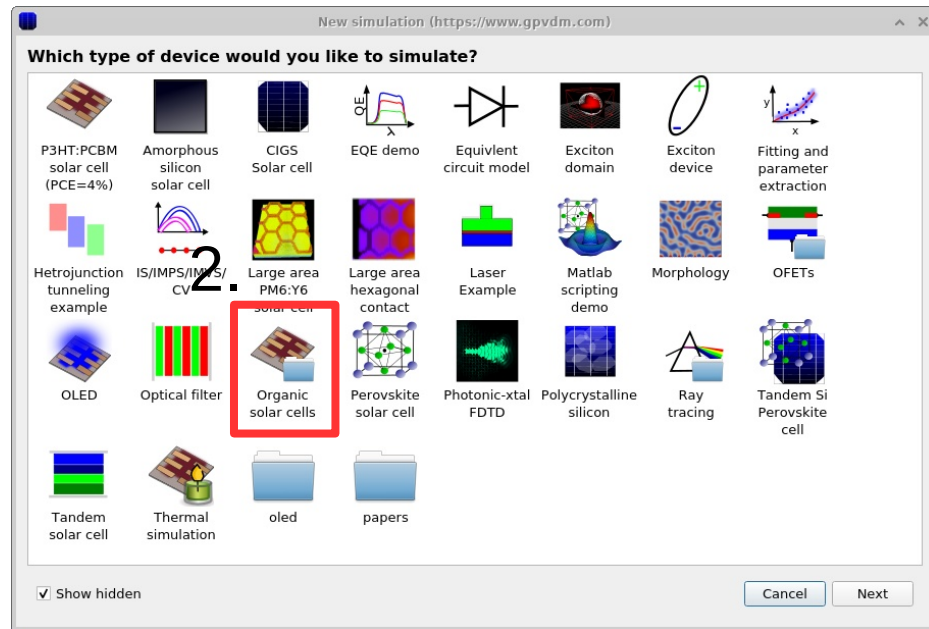
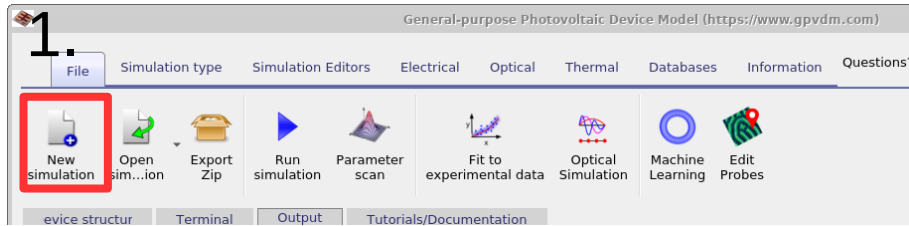
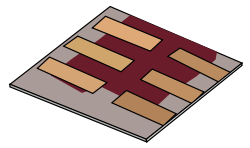
# We are now going to do an example using a solar cell.

- There is currently a lot of research interest in these devices.
- For this we will use the transfer matrix method. This assumes light propagates in the device as a wave.
- Don't worry about this too much for now but it's different to ray tracing.



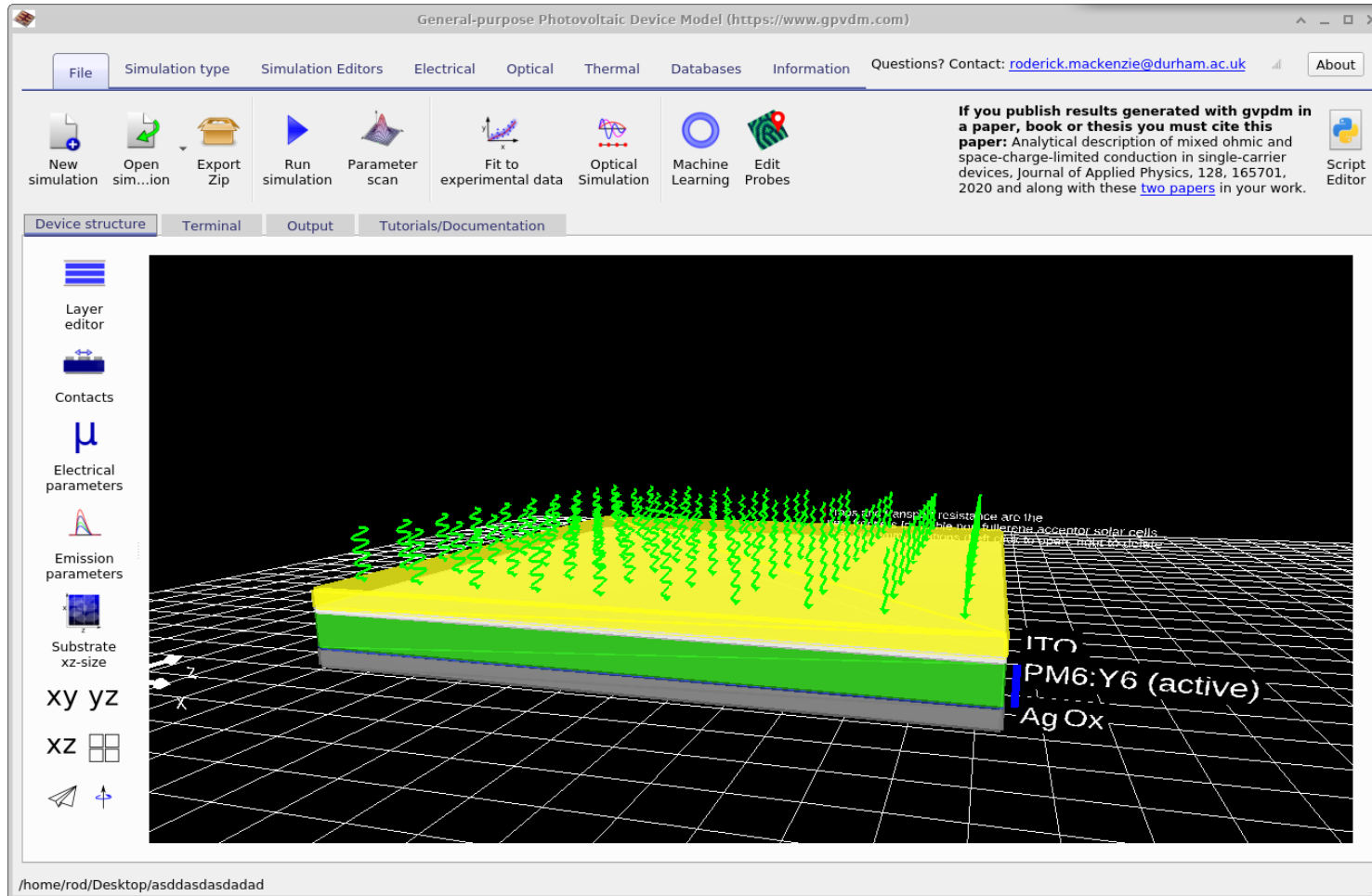
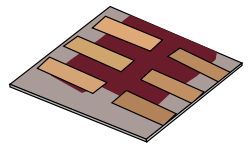
<https://www.nrel.gov/pv/organic-photovoltaic-solar-cells.html>

# Make a new solar cell simulation



- We are going to select a PM6:Y6 Organic solar cell which is a modern type of organic solar cell.
- Save this example to your home directory.

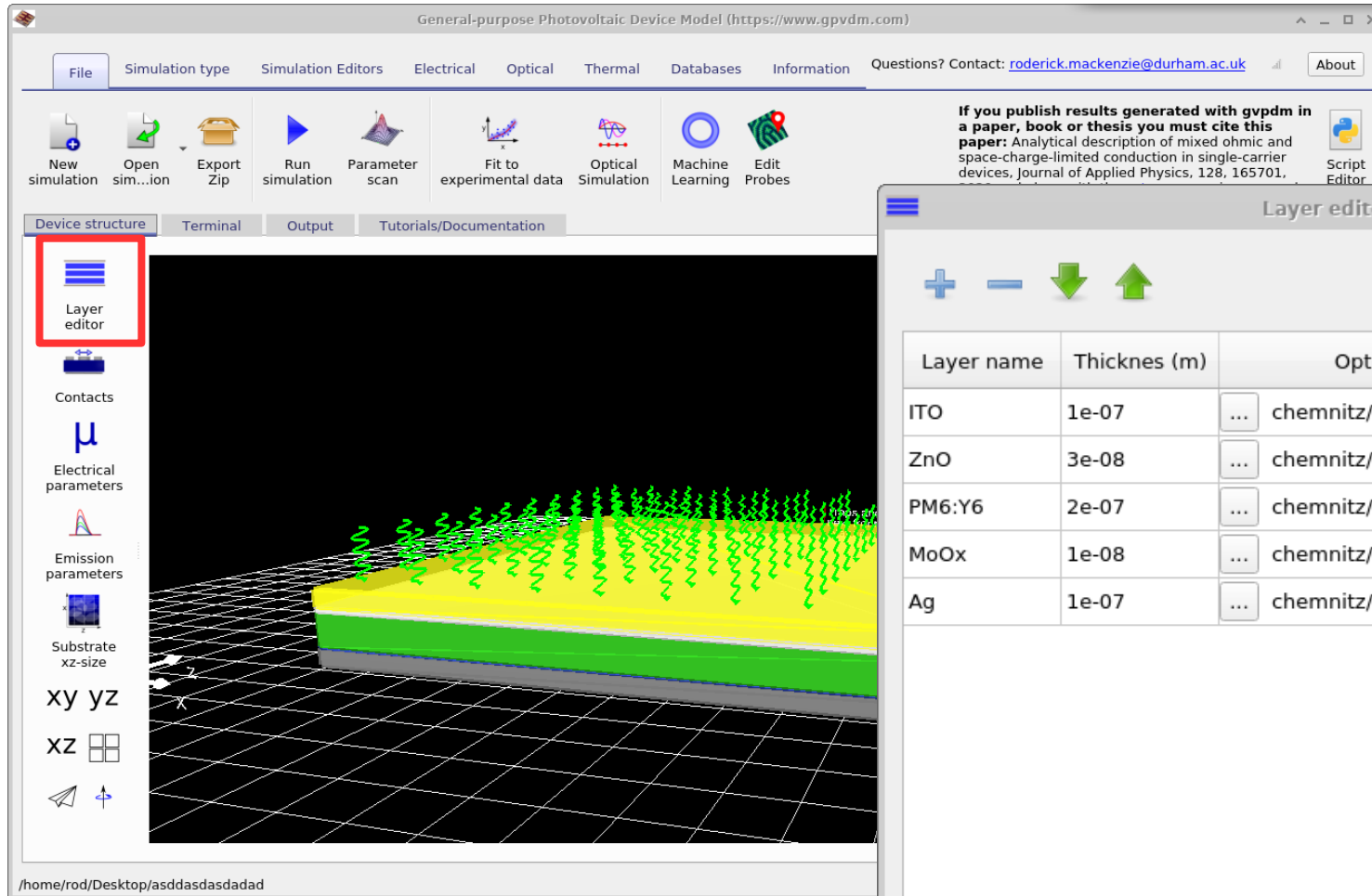
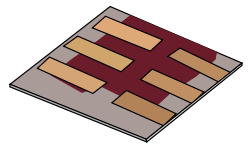
# You should get a window like this



- You can see the solar cell is made up of around five layers.
- You can see light coming in from the top.
- If you click on the layer editor you can inspect the layers.



# Opening the layer editor

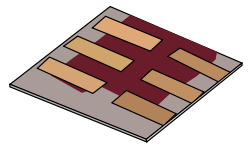


Layer editor <https://www.gvpdm.com>

Icons: +, -, ↓, ↑

Layer name	Thicknes (m)		Optical material	Layer type	Solve optical problem	Solve thermal problem	ID
ITO	1e-07	...	chemnitz/ito	contact	Yes	Yes	...
ZnO	3e-08	...	chemnitz/ZnO	other	Yes	Yes	...
PM6:Y6	2e-07	...	chemnitz/active_fresh	active layer	Yes	Yes	...
MoOx	1e-08	...	chemnitz/MoOx	other	Yes	Yes	...
Ag	1e-07	...	chemnitz/Ag	contact	Yes	Yes	...

# The layer editor



Layer editor <https://www.gpvdn.com>

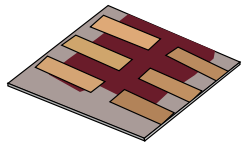
+ - ↓ ↑

Layer name	Thicknes (m)	Optical material	Layer type	Solve optical problem	Solve thermal problem	ID
ITO	1e-07	... chemnitz/ito	contact	Yes	Yes	...
ZnO	3e-08	... chemnitz/ZnO	other	Yes	Yes	...
PM6:Y6	2e-07	... chemnitz/active_fresh	active layer	Yes	Yes	...
MoOx	1e-08	... chemnitz/MoOx	other	Yes	Yes	...
Ag	1e-07	... chemnitz/Ag	contact	Yes	Yes	...

- **Layer name:** An English name for the layer, this has no technical significance (Tip: It might not like names with non English characters, i.e. Chinese characters)
- **Thickness of the layer:** The thickness of the layer in meters.
- **Optical material:** This points to the n/k data in the materials database. Use the “...” button to select a new material.
- **Other columns:** Discuss elsewhere.
- You can use the + button to add layers, the – button to remove layers.

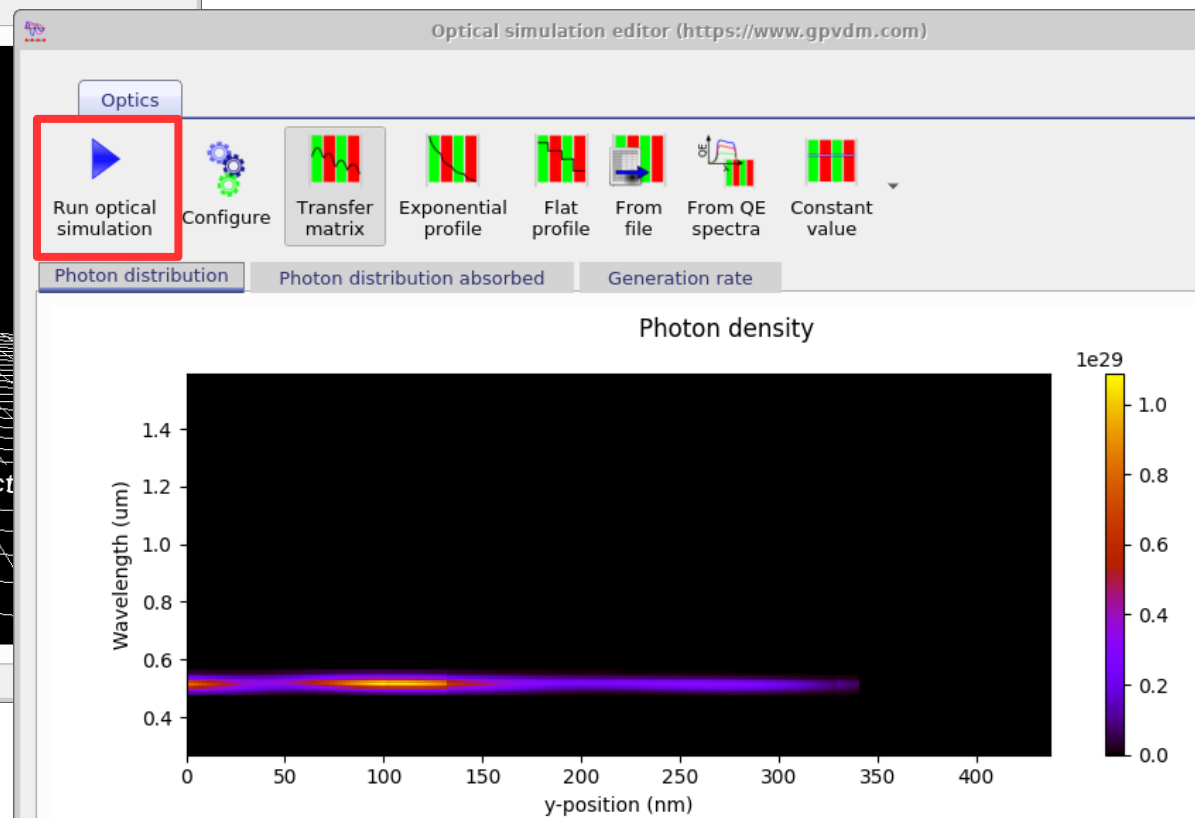
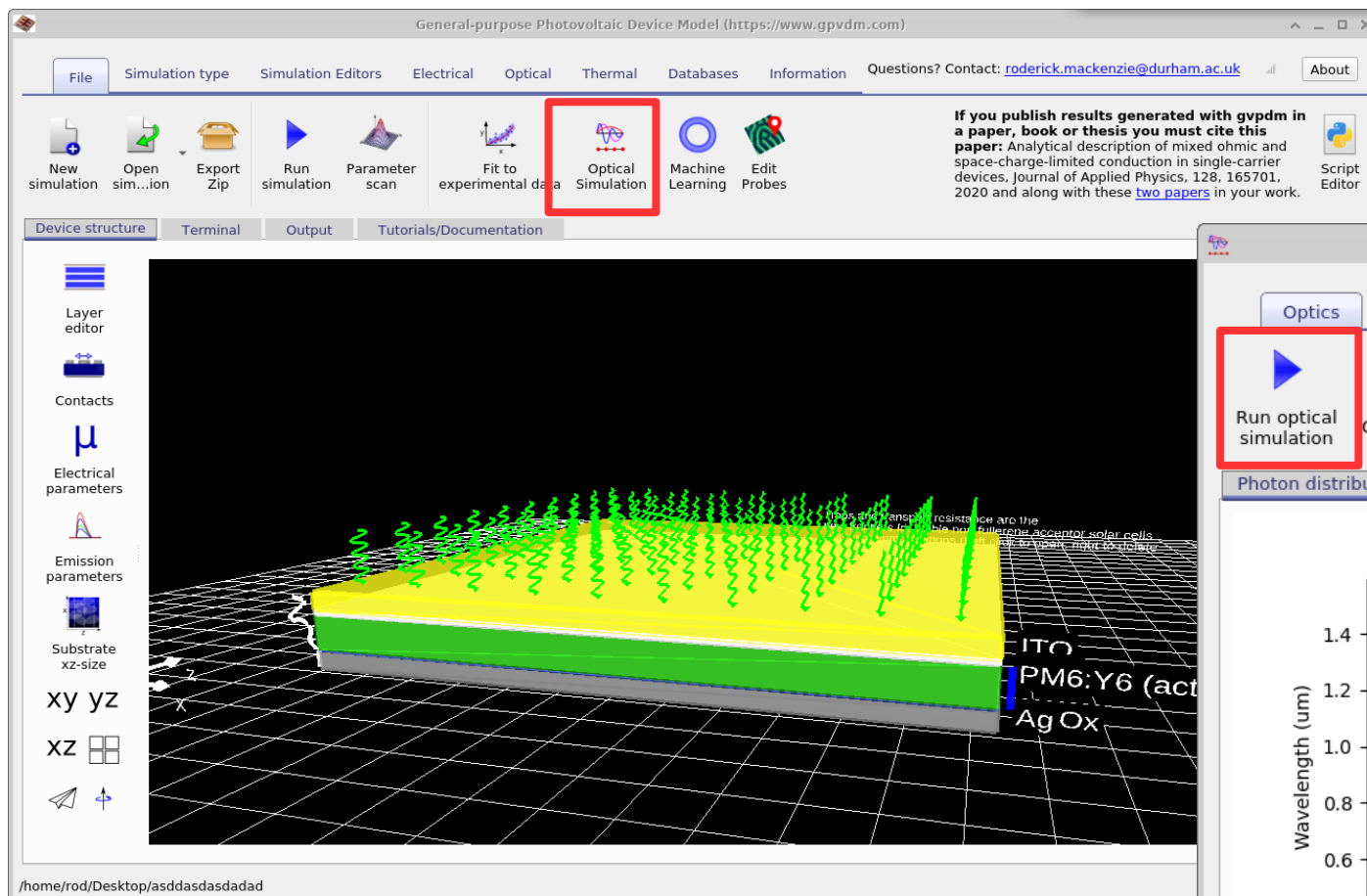
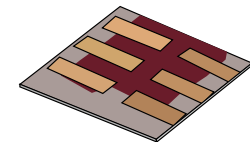


# Outline of the talk



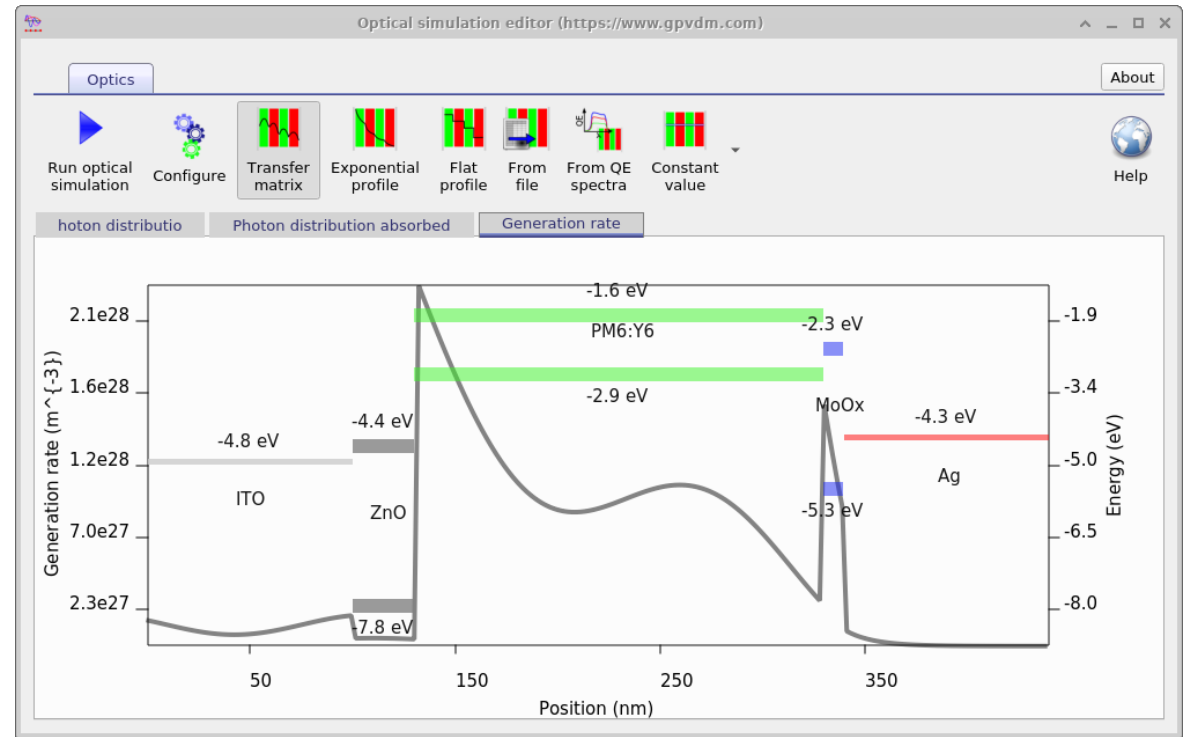
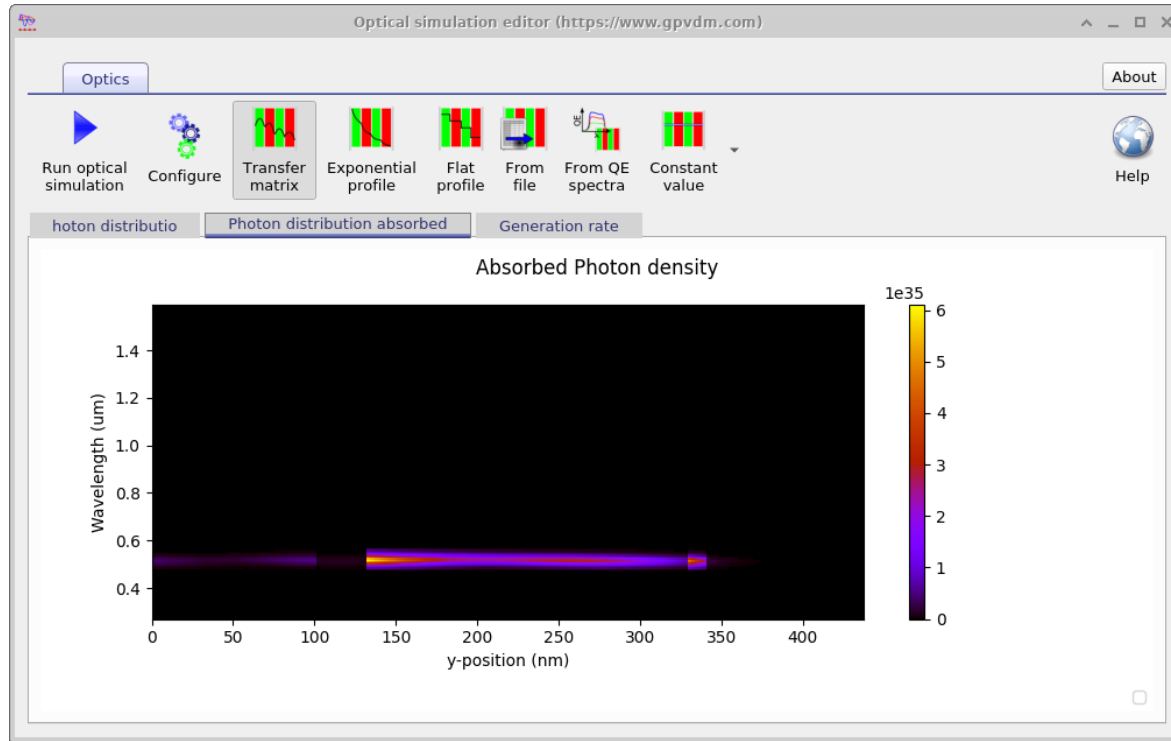
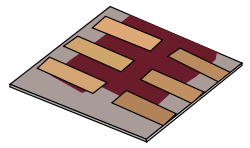
- In this talk we will cover:
  - What are optical simulations?
  - Why perform optical simulations?
  - What you need for accurate optical simulations
    - » Optical spectra
    - » Refractive index data ( $n$ )
    - » Optical absorption data ( $k$ )
      - The materials database
      - Importing  $n/k$  data into the model.
  - Setting up device structures
  - **Running optical simulations using OghmaNano.**
  - Light sources
  - Output files
  - Optical filter design task
  - Summary

# Running the full optical simulation...



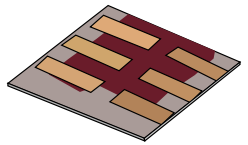
- Open the optical simulation window and click “Run optical simulation”

# Look at the generation rate in the device:



- This cell is excited with an LED, so you will see a very narrow region of excitation at 515 nm.
- The generation rate tab shows the generation profile in the device with all wavelengths jointed together.

## Exercise 2: Play with the layer thickness and see how it influences the distribution of light in the device.



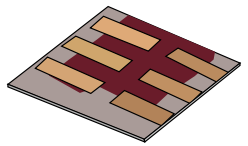
Layer editor <https://www.gpvdn.com>

+ - ↓ ↑

Layer name	Thicknes (m)	Optical material	Layer type	Solve optical problem	Solve thermal problem	ID
ITO	1e-07	... chemnitz/ito	contact	Yes	Yes	...
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PM6:Y6	2e-07	... chemnitz/active_fresh	active layer	Yes	Yes	...
MoOx	1e-08	... chemnitz/MoOx	other	Yes	Yes	...
Ag	1e-07	... chemnitz/Ag	contact	Yes	Yes	...

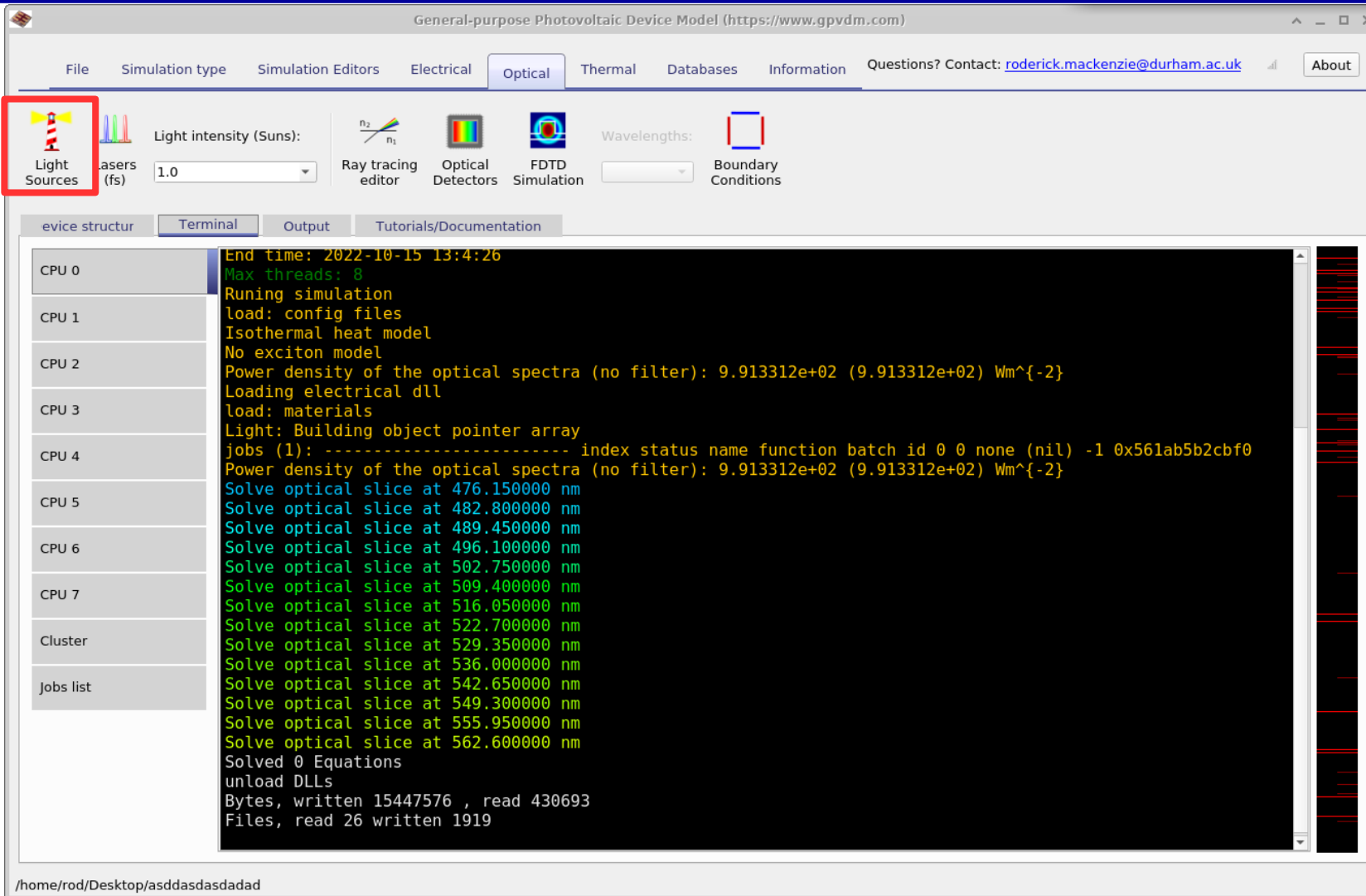
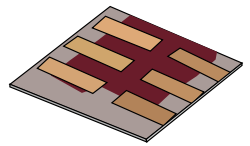
- Increase the thickness of the active layer to 300nm and change the material to chemnitz/active\_aged.

# Outline of the talk



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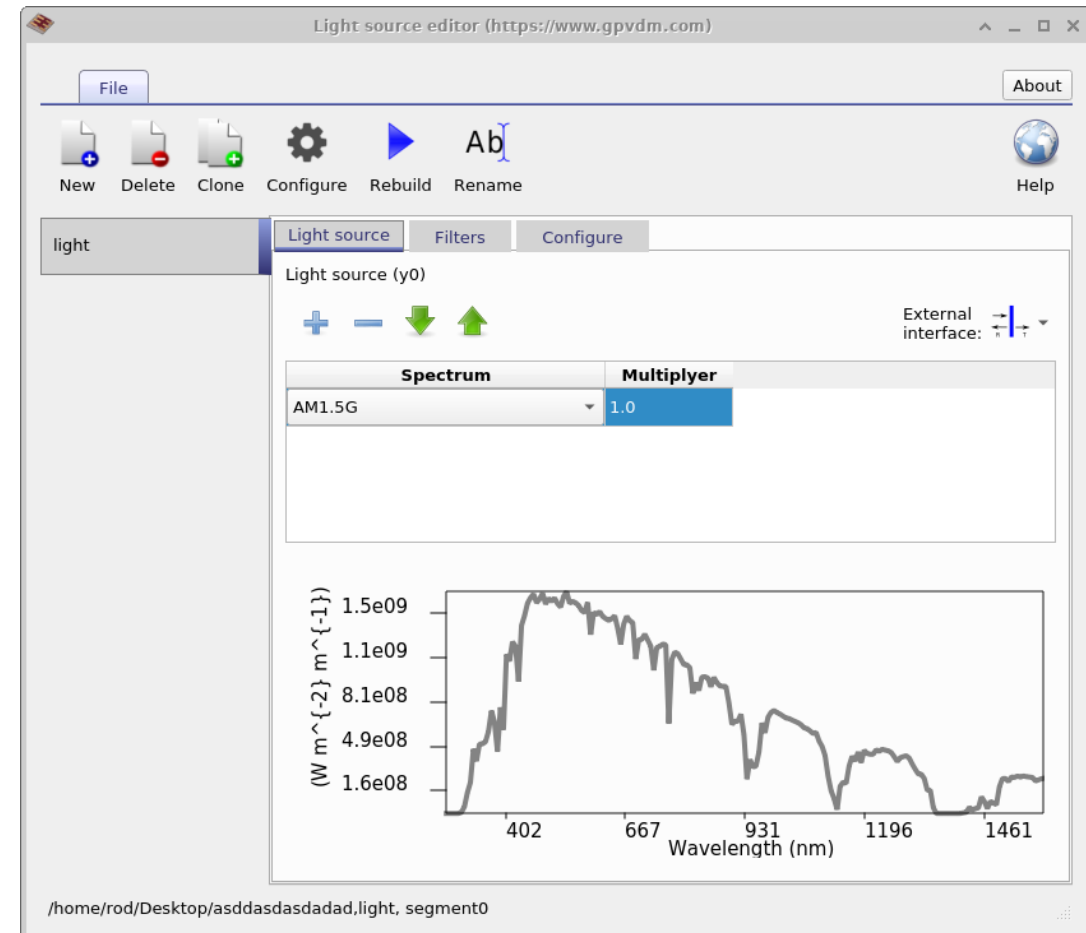
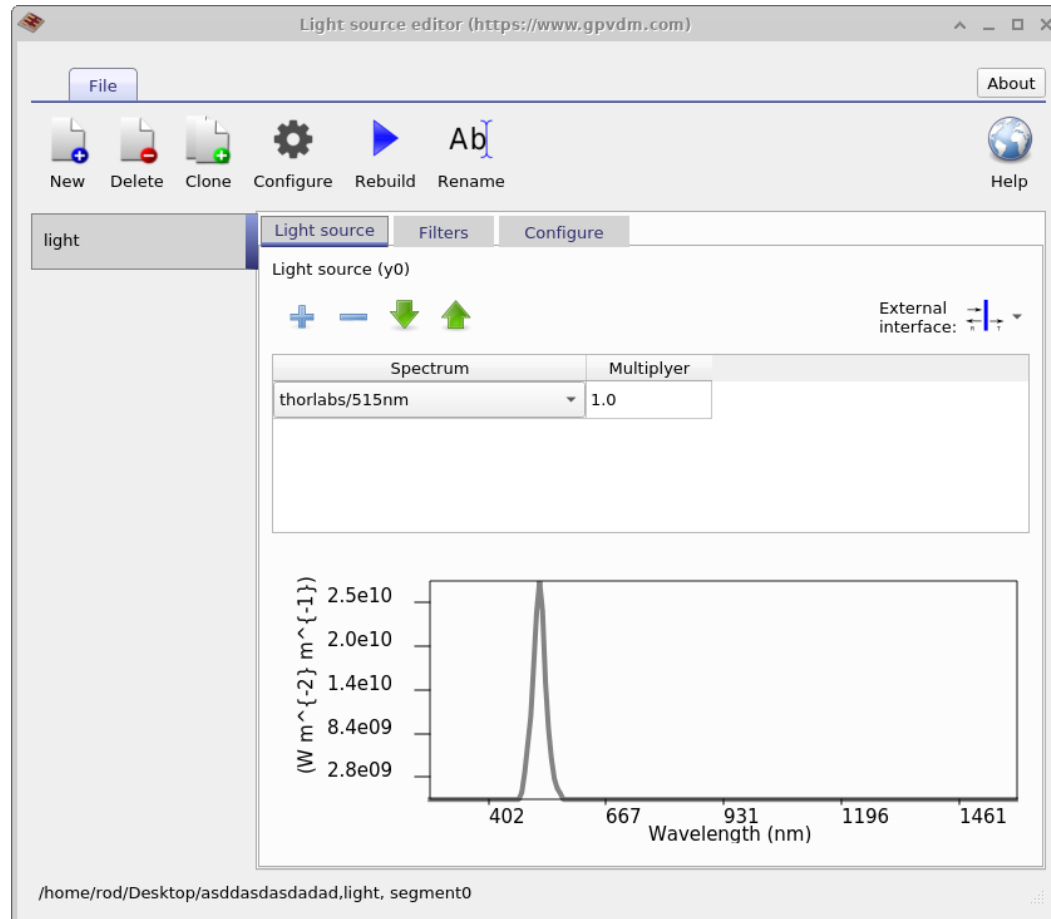
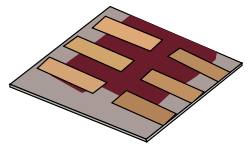
# Selecting light sources



- You will have noticed that the light used to illuminate the last simulation was an LED with a very narrow spectrum.

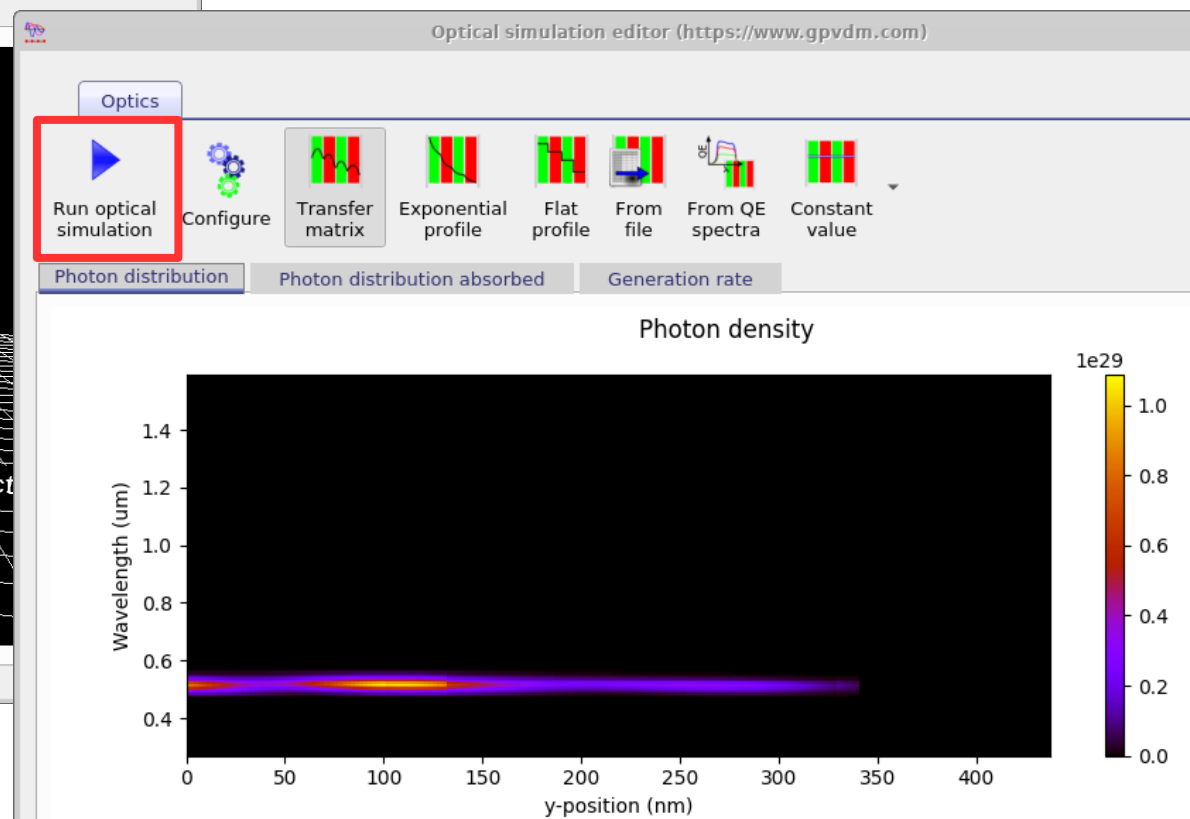
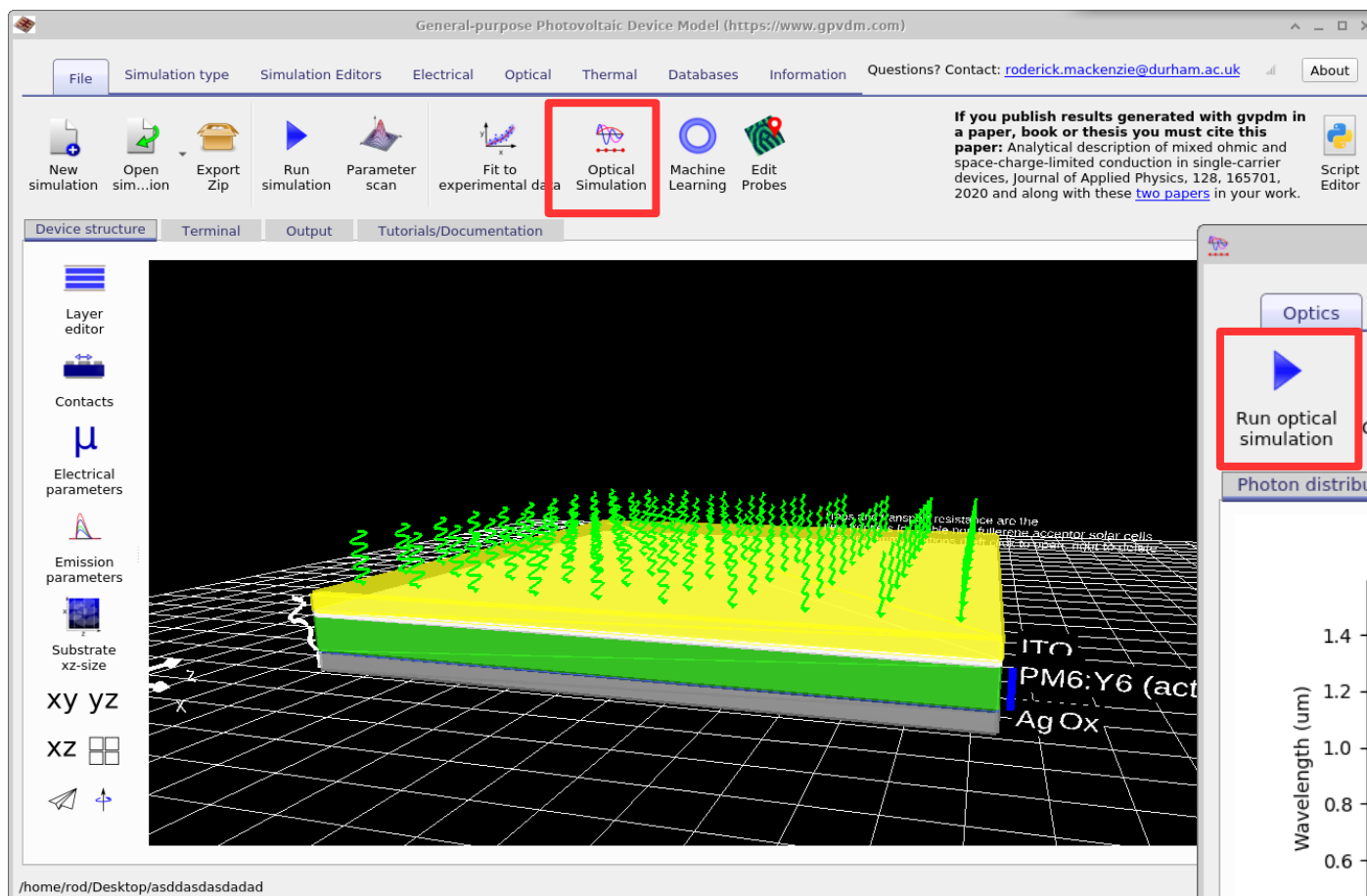
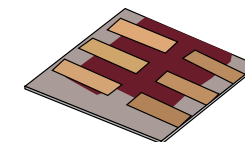
- We can change this to something more realistic by clicking on the Light sources button in the optical ribbon.

# Selecting a Light source:



•**Note:** You can also mix various spectra and apply filters using this window, try using the add button to mix the AM1.5G spectra and the 515nm laser from ThorLabs.

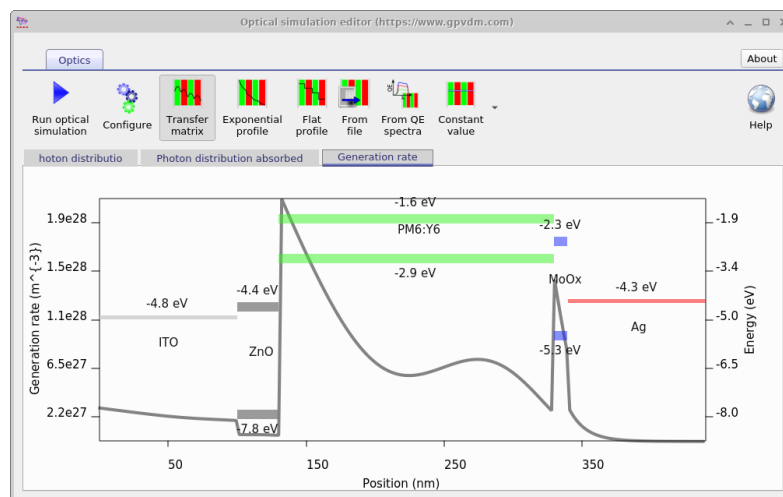
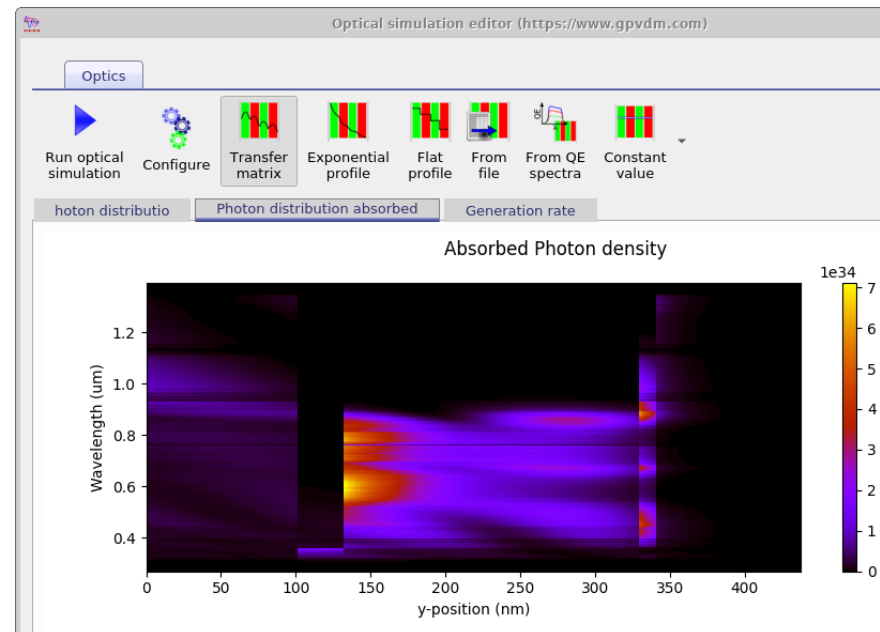
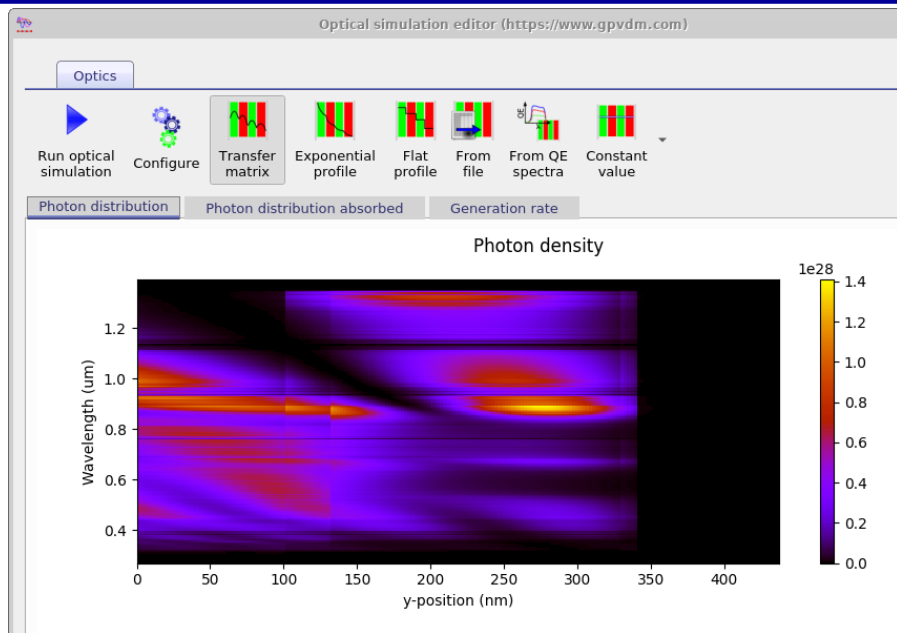
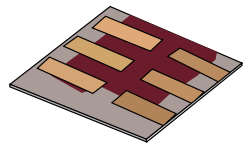
# Now go back and rerun the optical simulation



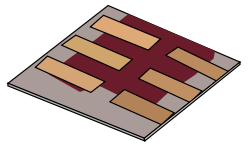
- Open the optical simulation window and click “Run optical simulation”



# Now go back and rerun the optical simulation

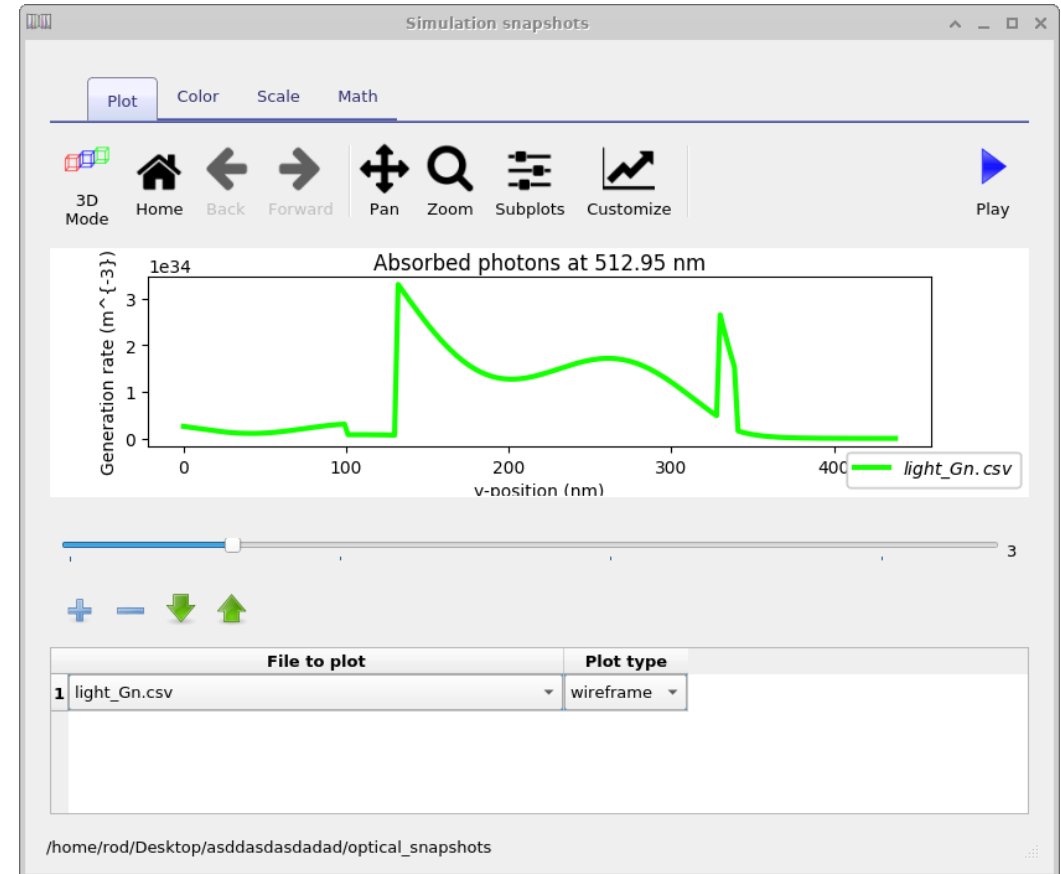
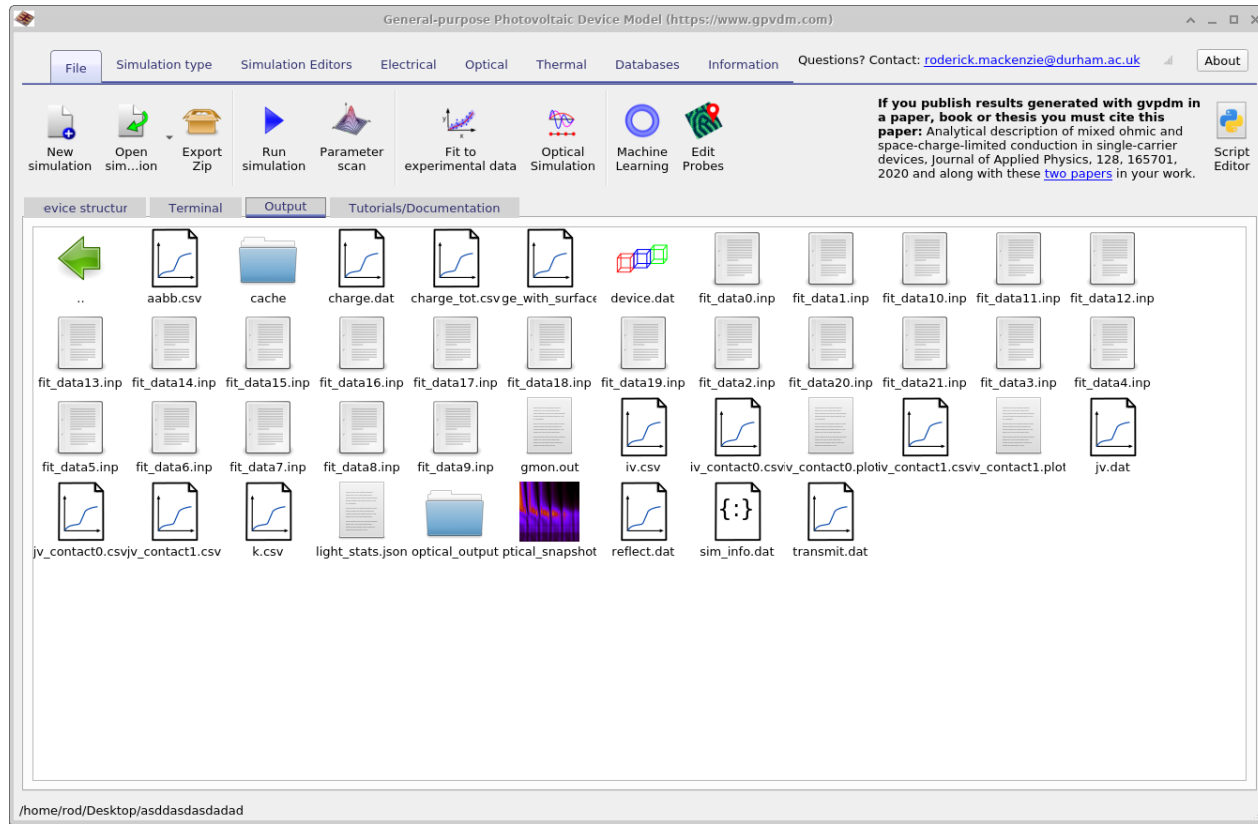
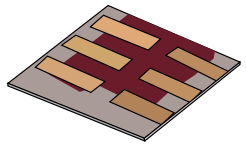


# Outline of the talk

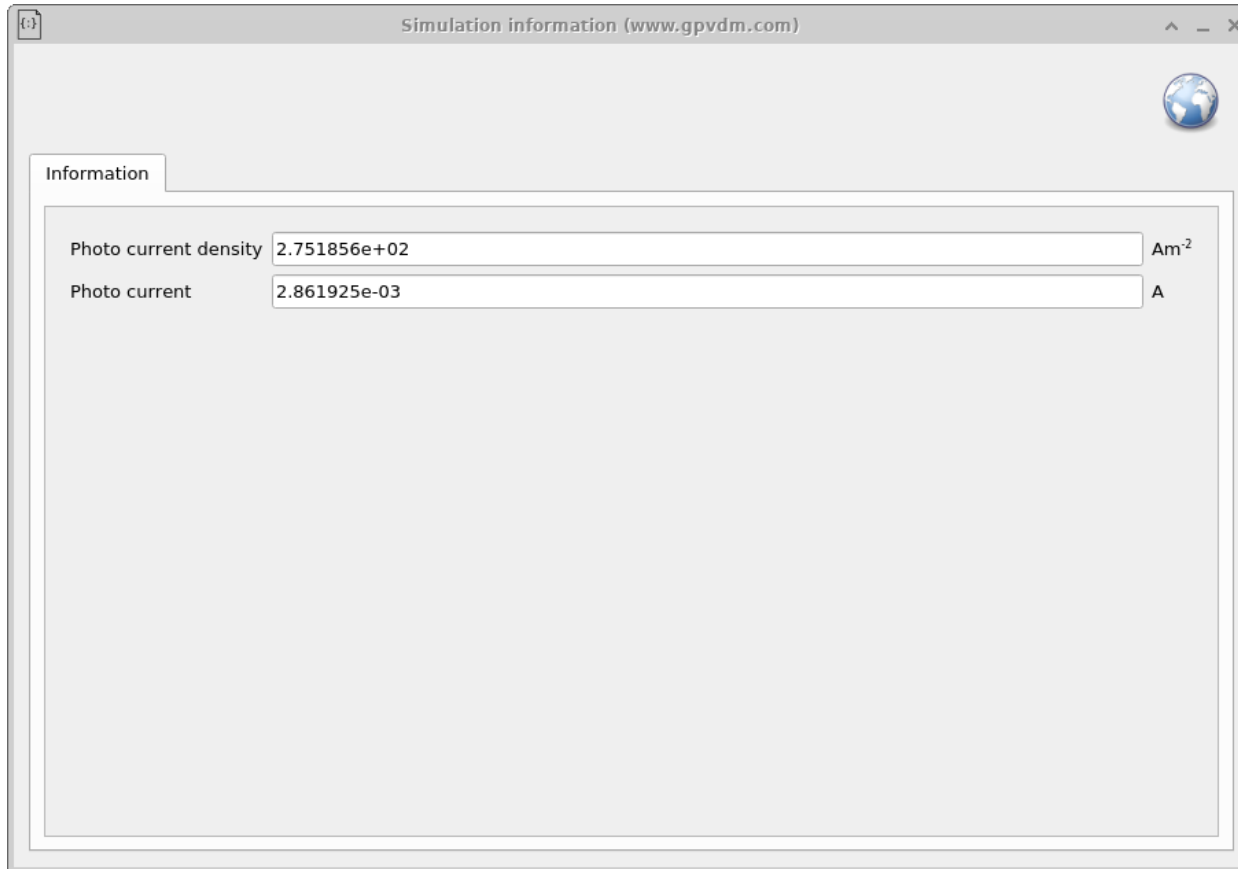
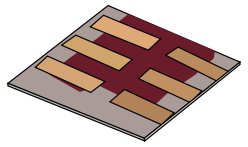


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# Outputs: Optical snapshots

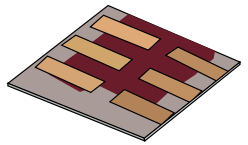


# Simulation information: sim\_info.dat



- This file contains the maximum photocurrent one would get out of the device.

# Statistics: light\_stats.json

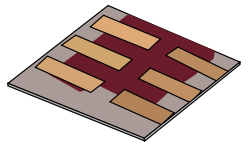


```
light_stats.json (~/Desktop/asddasdasdadad) - Pluma
File Edit View Search Tools Documents Help
Open Save Undo
light_stats.json x
1 {
2   "layers" : 5,
3   "layer0": {
4     "light_frac_photon_generation": 1.103641e-01
5   },
6   "layer1": {
7     "light_frac_photon_generation": 8.650944e-03
8   },
9   "layer2": {
10    "light_frac_photon_generation": 8.050166e-01
11  },
12  "layer3": {
13    "light_frac_photon_generation": 5.804029e-02
14  },
15  "layer4": {
16    "light_frac_photon_generation": 1.792815e-02
17  }
18 }
```

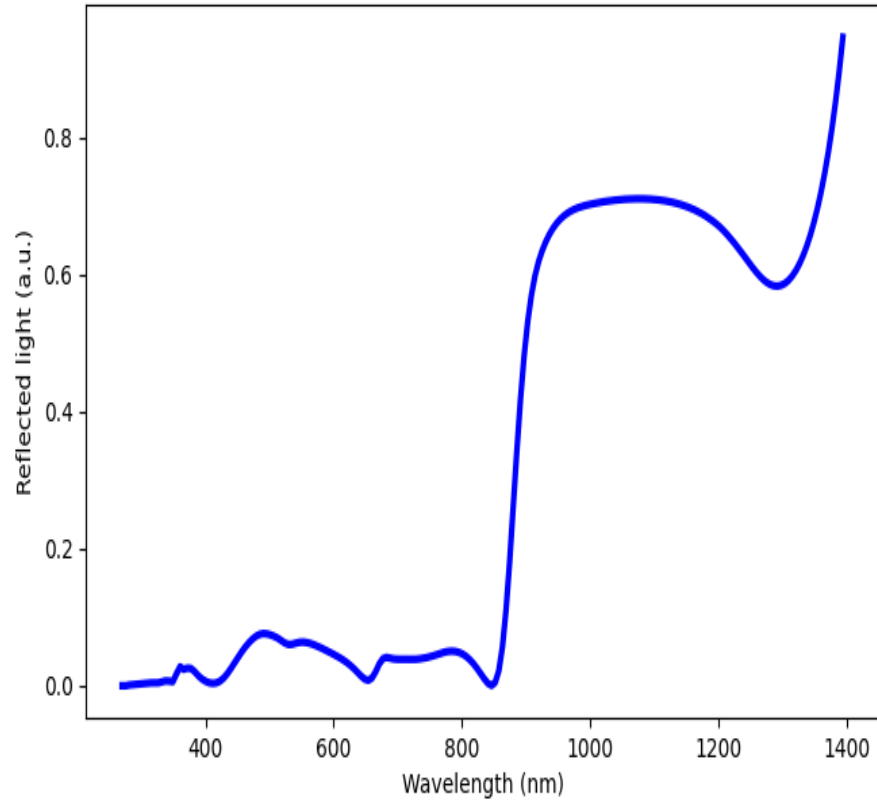
- This file contains a breakdown of which layers absorb what fraction of light in the material.

# Optical transmission/reflection

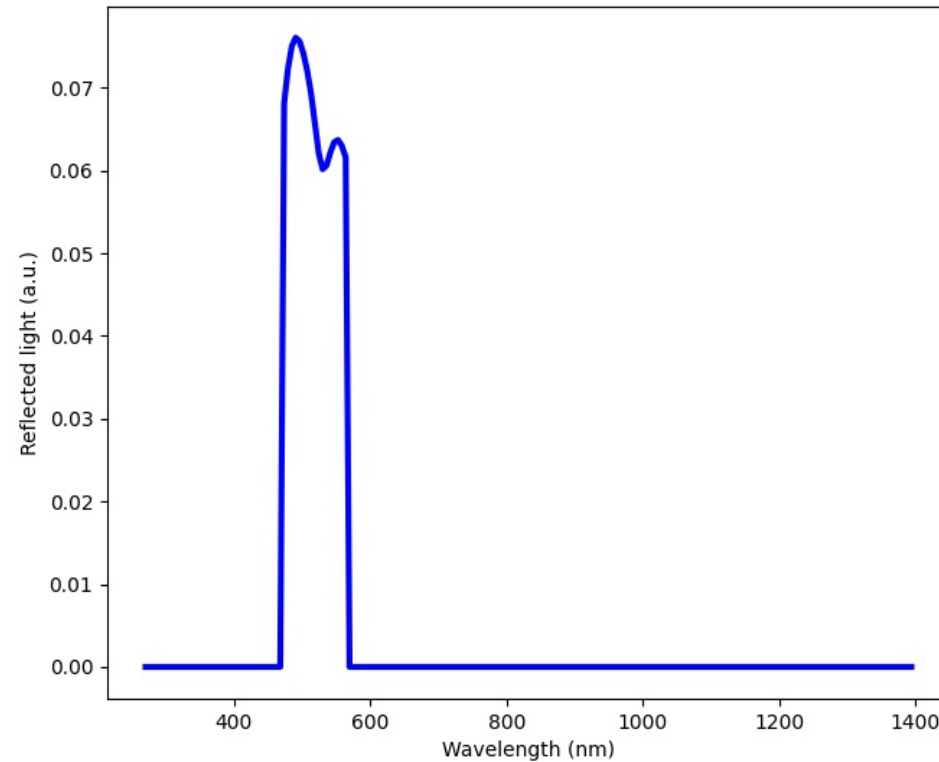
transmit.csv, reflect.csv



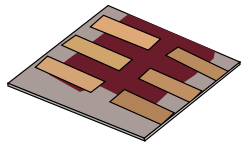
- Transmission (transmit.csv)



- Reflection (reflect.csv)

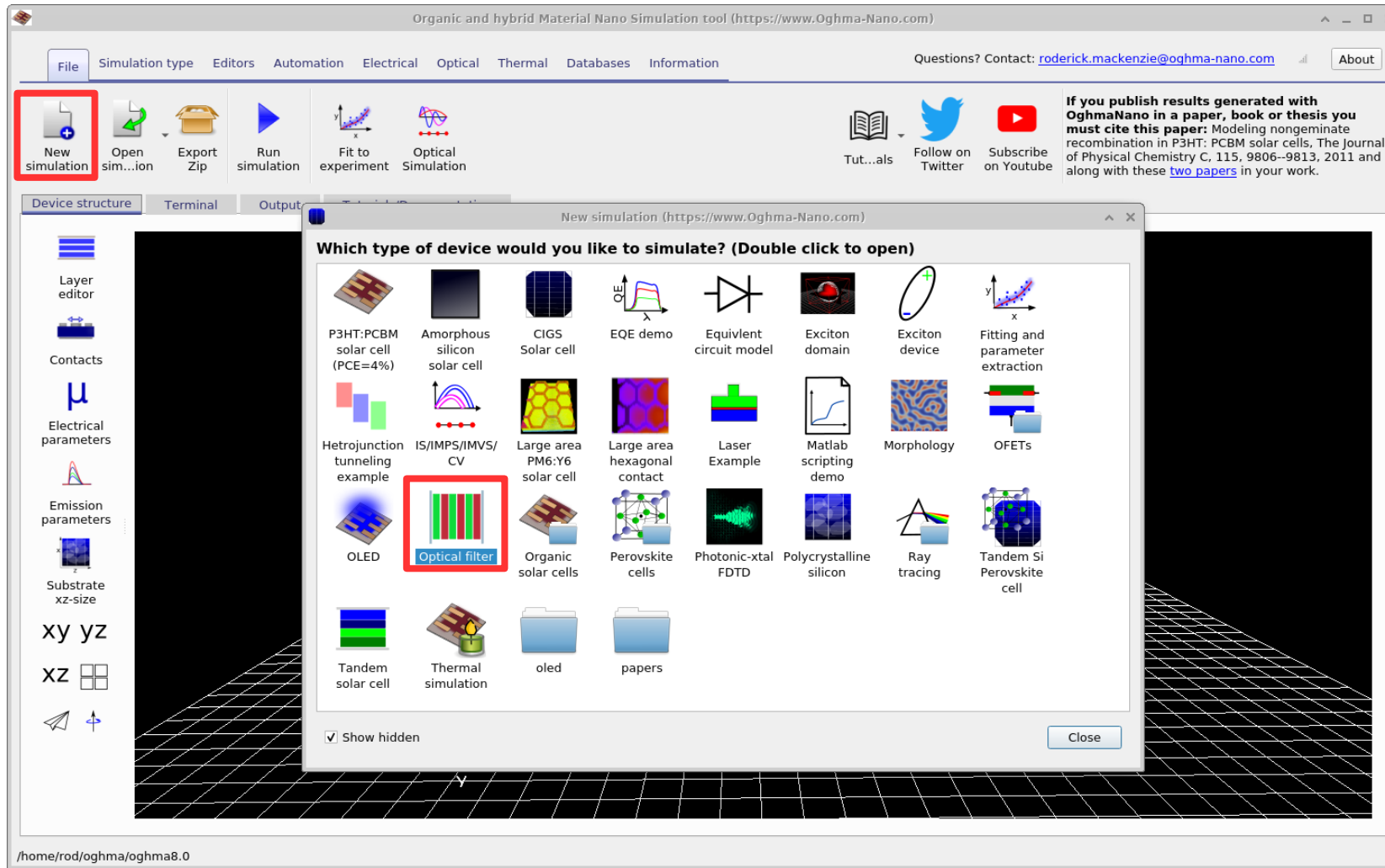
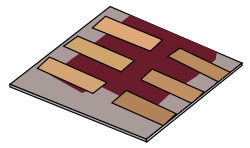


# Outline of the talk



- In this talk we will cover:
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# Task: Designing optical filters

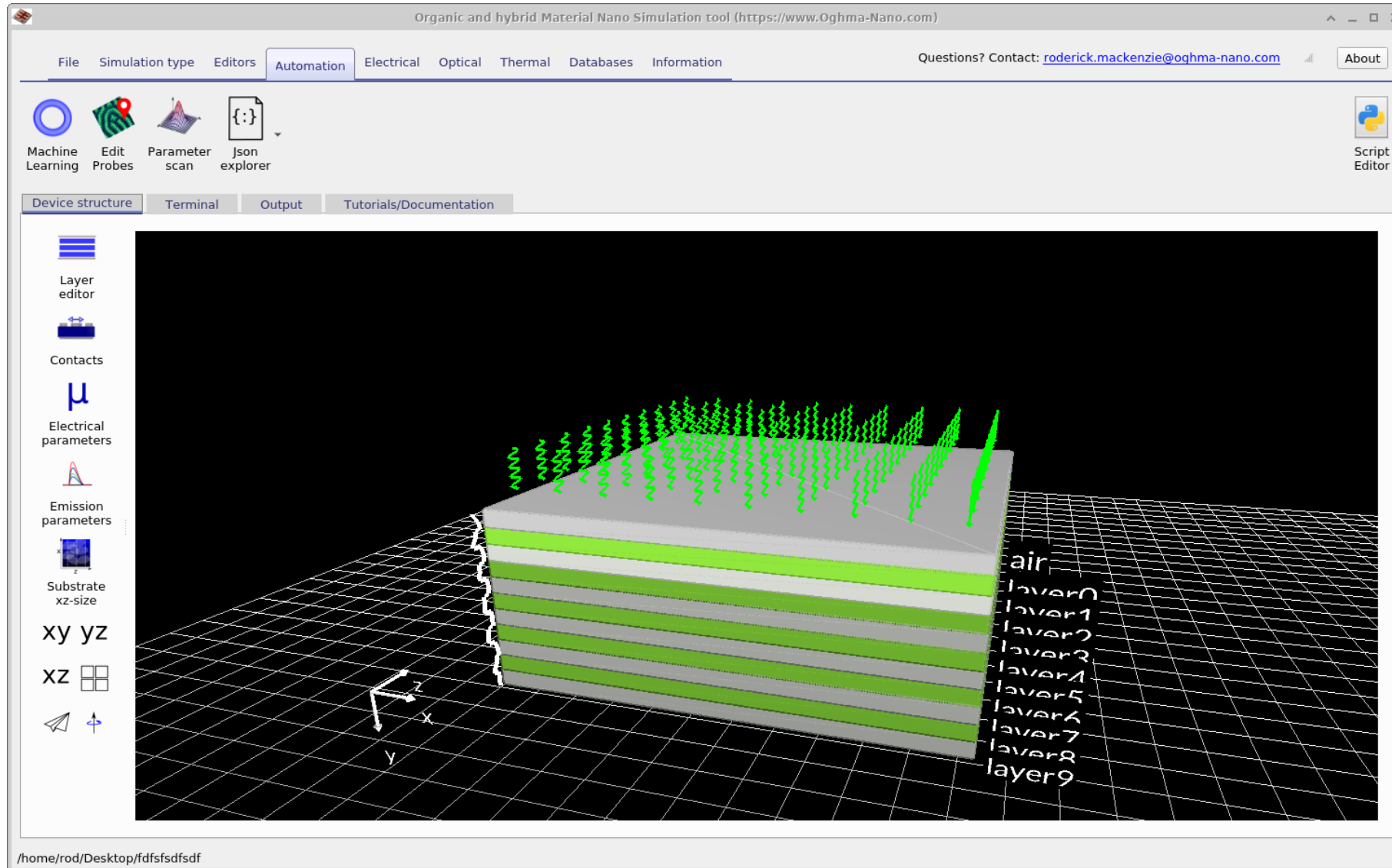
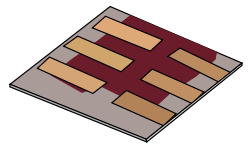


- First make an optical filter simulation.

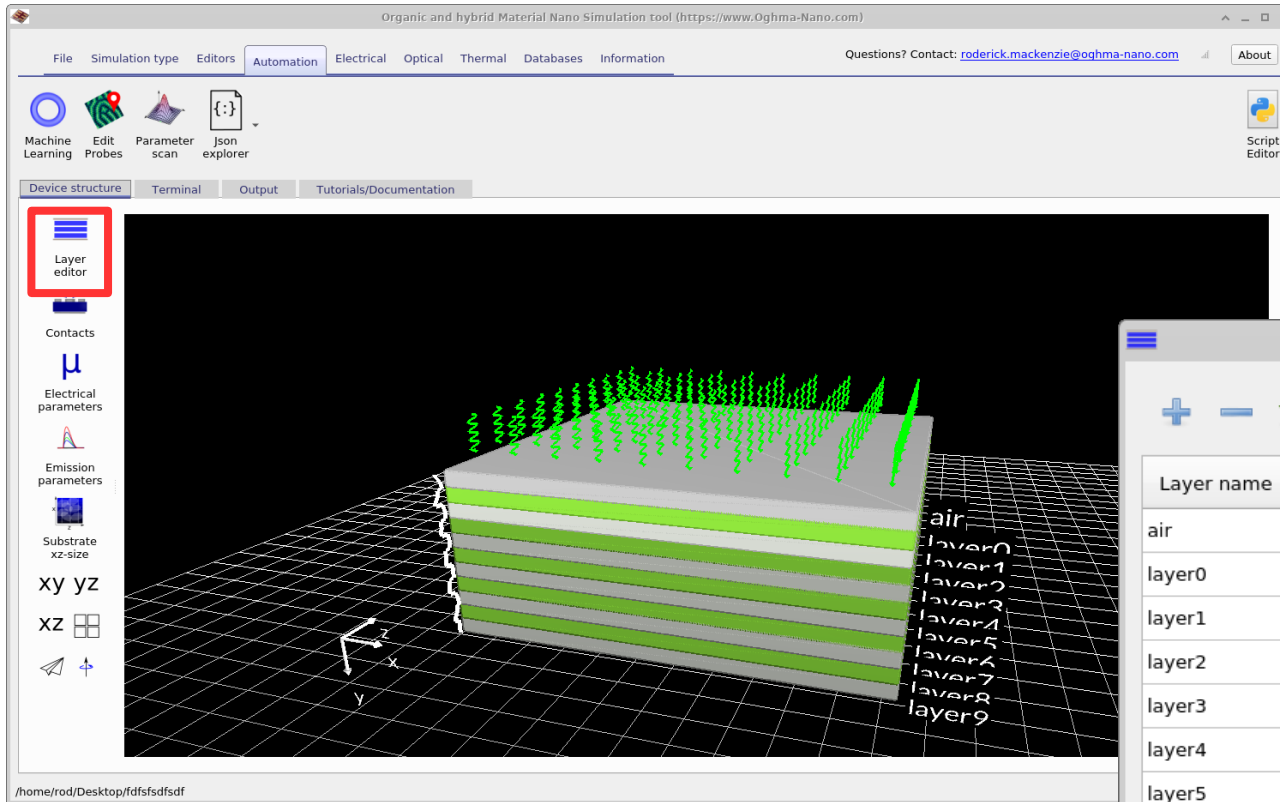
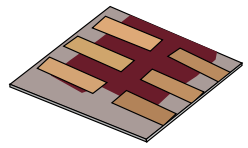
- Then run the simulation and examine the device structure, the reflective light and the transmitted light.



# It should bring up this window



# Open the layer editor to inspect the layers

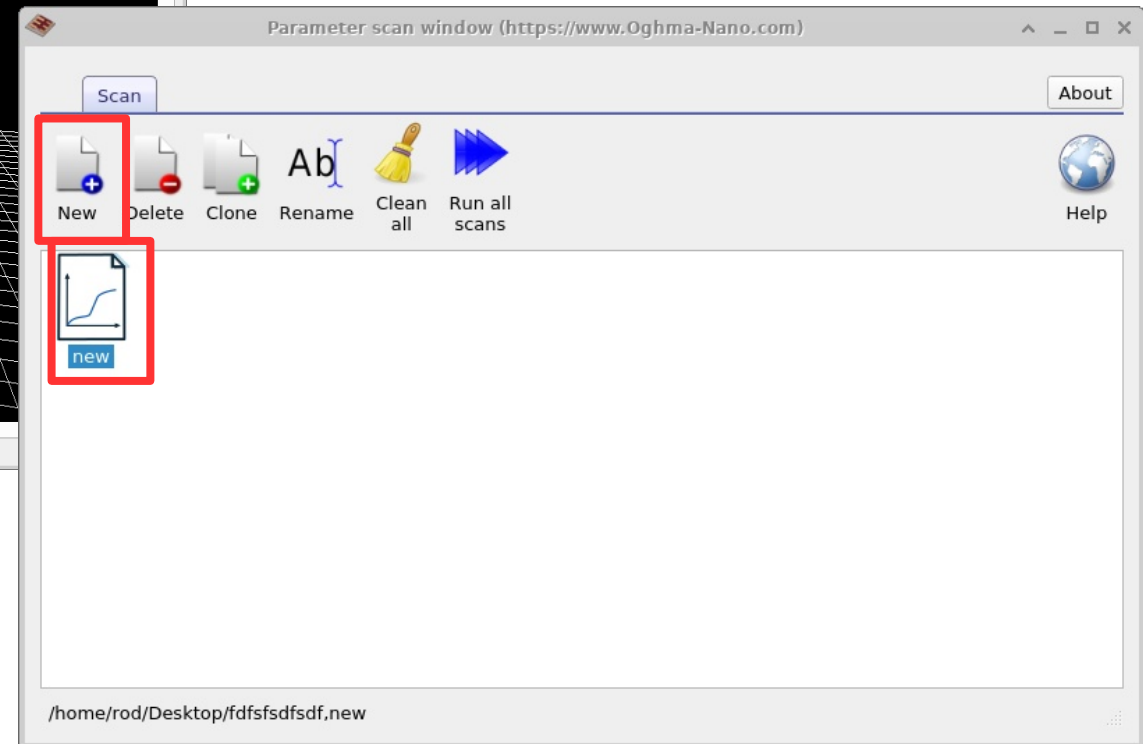
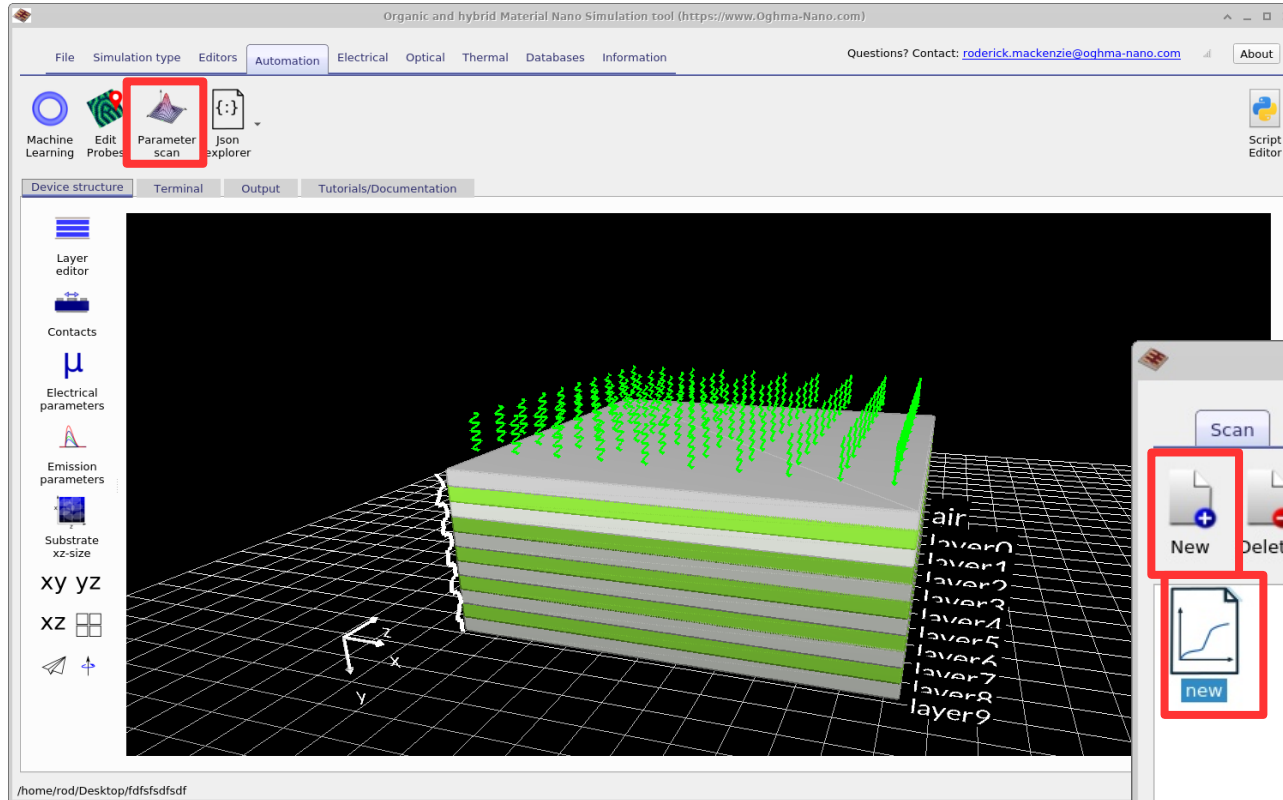
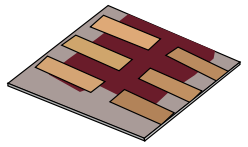


Layer editor (<https://www.Oghma-Nano.com>)

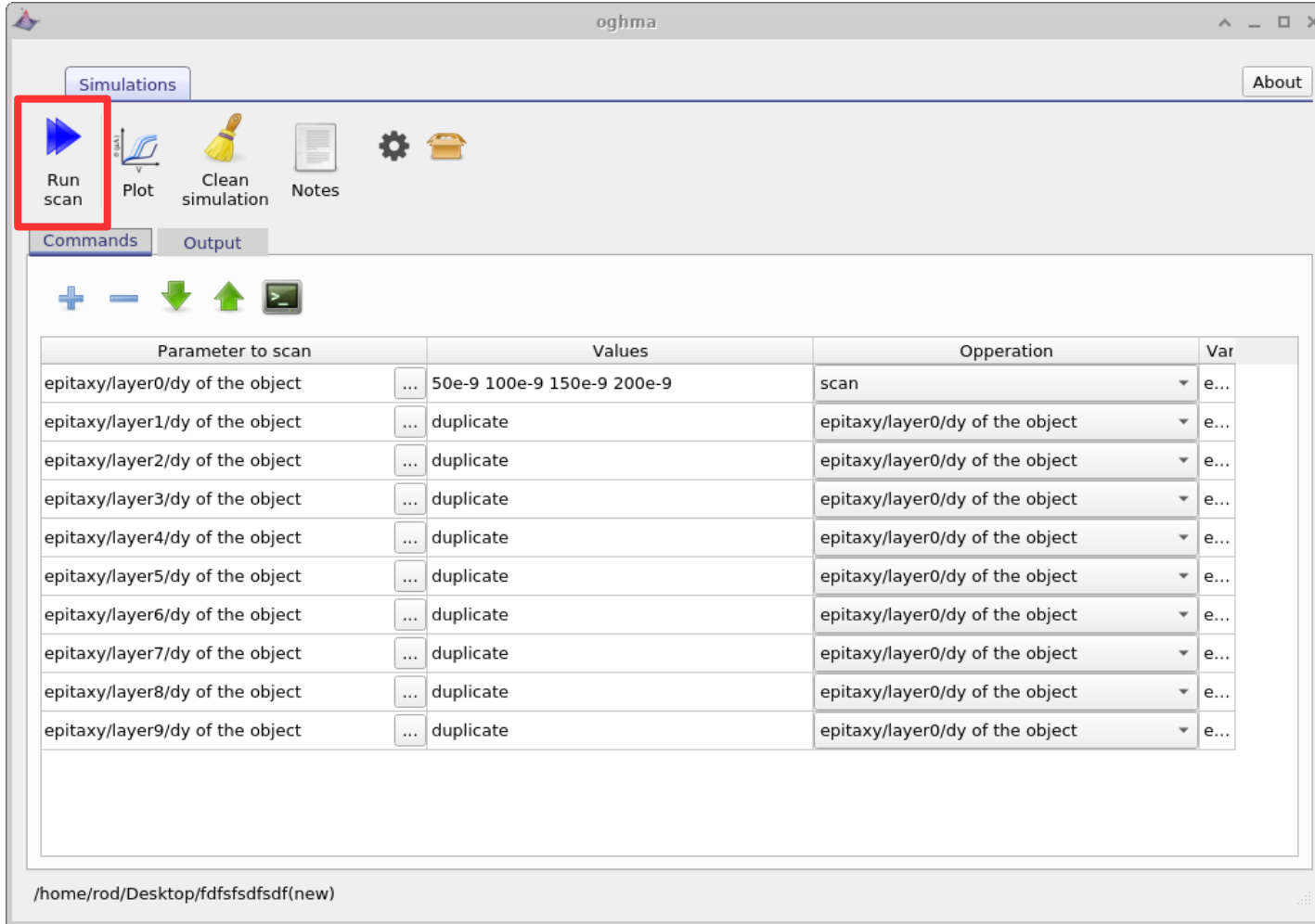
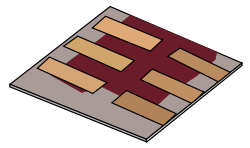
+ - ↓ ↑

Layer name	Thicknes (m)		Optical material	Layer type	Solve optical problem	Solve thermal problem	ID
air	1e-07	...	generic/air	other	Yes	Yes	i...
layer0	1e-07	...	generic/n_with_alpha/2.0	other	Yes	Yes	i...
layer1	1e-07	...	generic/n_with_alpha/1.0	other	Yes	Yes	i...
layer2	1e-07	...	generic/n_with_alpha/2.0	other	Yes	Yes	i...
layer3	1e-07	...	generic/n_with_alpha/1.0	other	Yes	Yes	i...
layer4	1e-07	...	generic/n_with_alpha/2.0	other	Yes	Yes	i...
layer5	1e-07	...	generic/n_with_alpha/1.0	other	Yes	Yes	i...
layer6	1e-07	...	generic/n_with_alpha/2.0	other	Yes	Yes	i...
layer7	1e-07	...	generic/n_with_alpha/1.0	other	Yes	Yes	i...
layer8	1e-07	...	generic/n_with_alpha/2.0	other	Yes	Yes	i...
layer9	1e-07	...	generic/n_with_alpha/1.0	other	Yes	Yes	i...

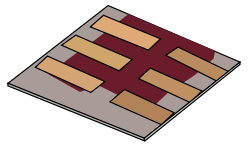
# Click on the parameter scan window,



# Then make the scan window look like this: And run the simulation



# After it has finished



oghma

Simulations

Run scan Plot Clean simulation Notes

Commands Output

Parameter to scan	Values	Operation	Var
epitaxy/layer0/dy of the object	50e-9 100e-9 150e-9 200e-9	scan	e...
epitaxy/layer1/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer2/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer3/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer4/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer5/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer6/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer7/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer8/dy of the object	...	duplicate	epitaxy/layer0/dy of the object
epitaxy/layer9/dy of the object	...	duplicate	epitaxy/layer0/dy of the object

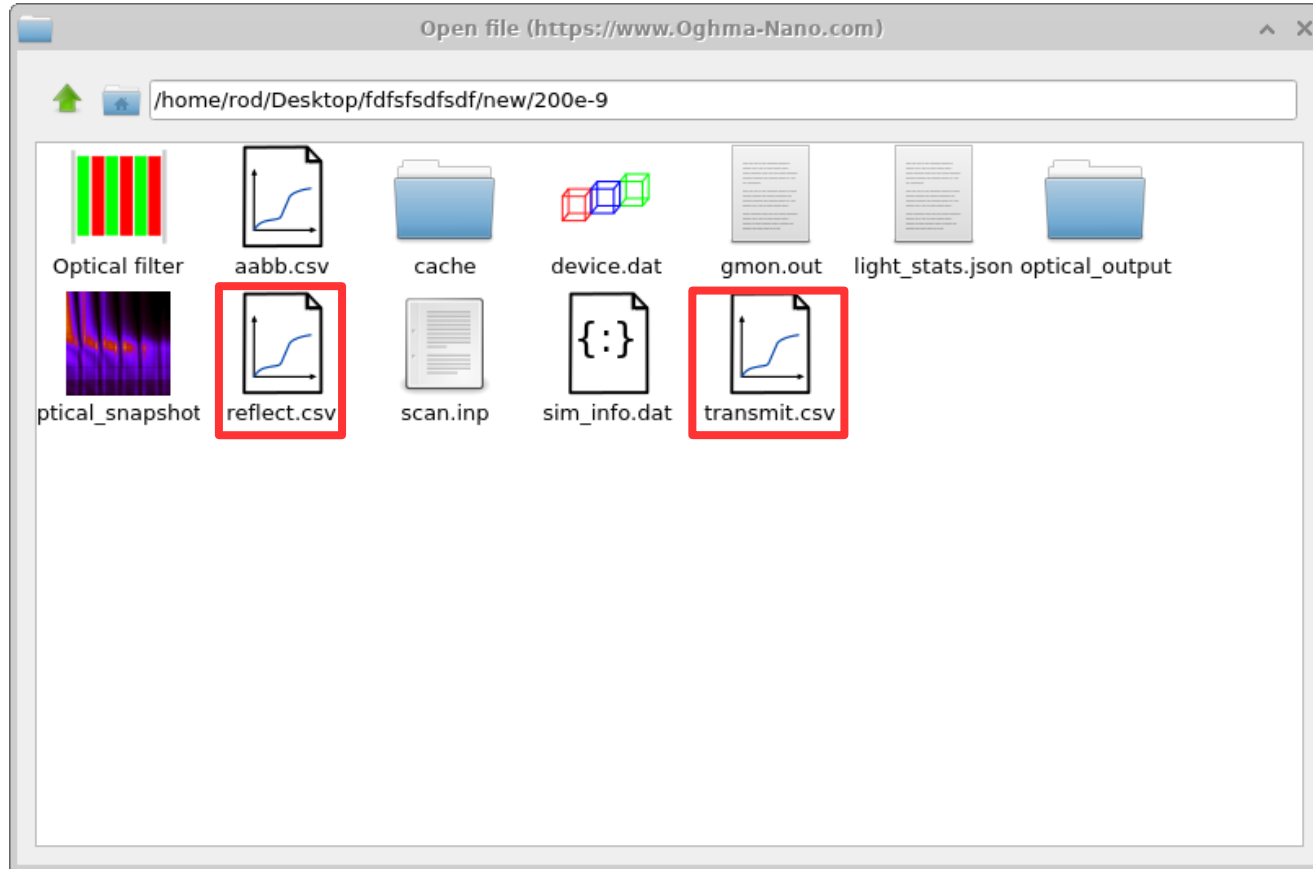
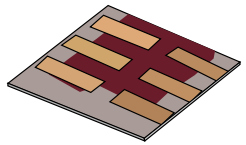
/home/rod/Desktop/fdfsfsdfsdf(new)

Open file (https://www.Oghma-Nano.com)

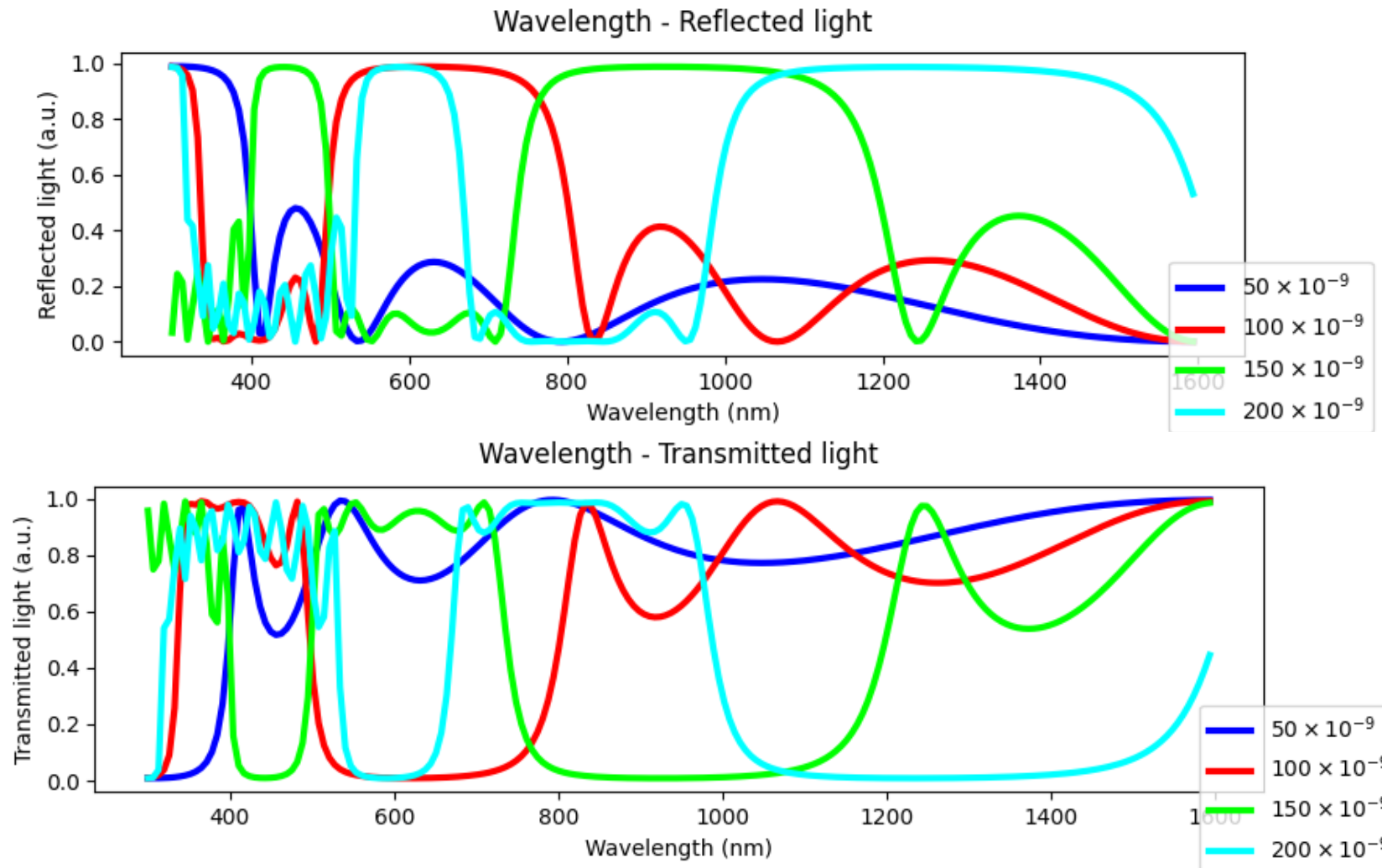
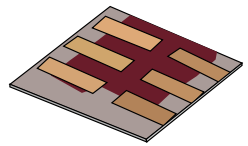
/home/rod/Desktop/fdfsfsdfsdf/new

100e-9 150e-9 200e-9 50e-9 flat\_list.inp

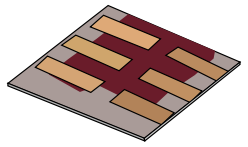
# Then examine these files



They should look like this:



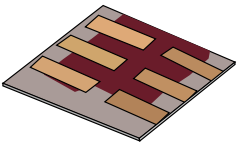
## Now your go:



- By changing the refractive index and the layer thicknesses try to make the filter transmit light between 500-700 nm and reflect light at 500-700 nm .



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  - Running optical simulations using gpvdm
  - Light sources
  - Output
  - **Summary**